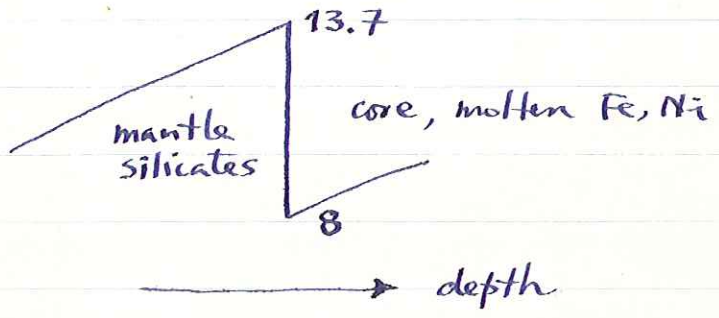


Before considering the inverse problem we must pause to consider the complications which may arise in $T(\Delta)$ and $\beta(\Delta)$ curves.

First, the most important features caused by the fluid core.

The P velocity decreases from ~ 13.7 km/s to ~ 8 km/s at core-mantle bdry

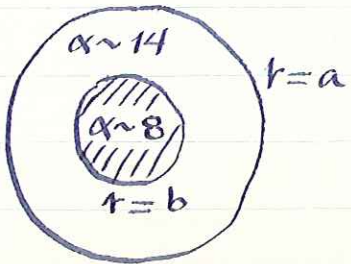


Main reason: fluid $\mu \rightarrow 0$, $\alpha = \left(\frac{K + \frac{4}{3}\mu}{\rho} \right)^{1/2}$

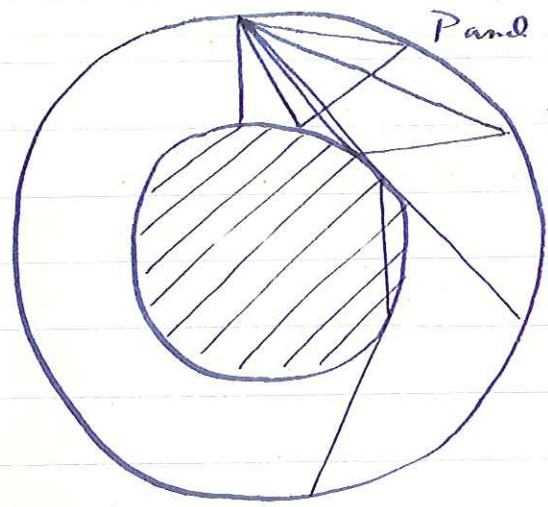
↑
not only reason, also major compositional difference.

$\alpha_{core} < \alpha_{mantle}$
 $\alpha_{core} > \beta_{mantle}$, but $\approx \beta_{mantle} = 7.2$ km/s

Consider simplest example: homogeneous core and mantle, e.g.

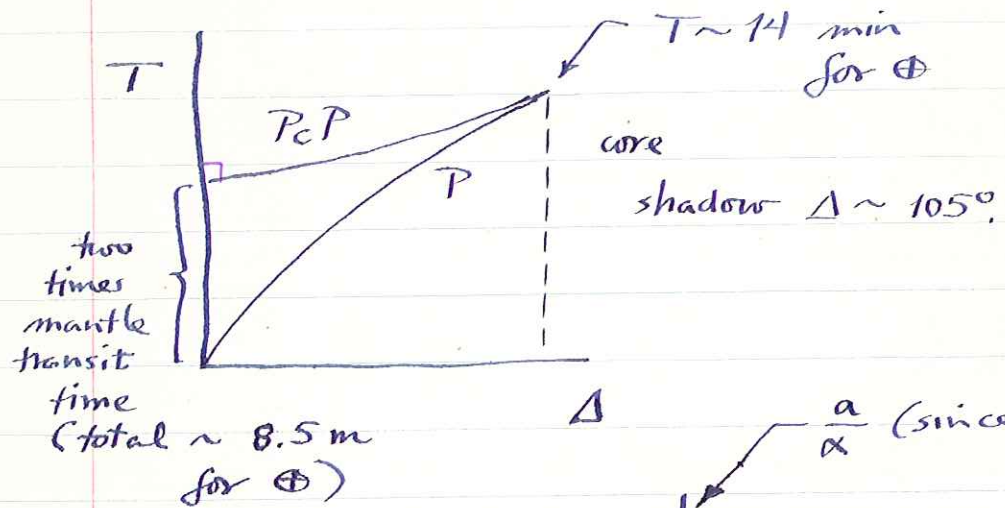


Surface focus quake: consider first P and PcP.



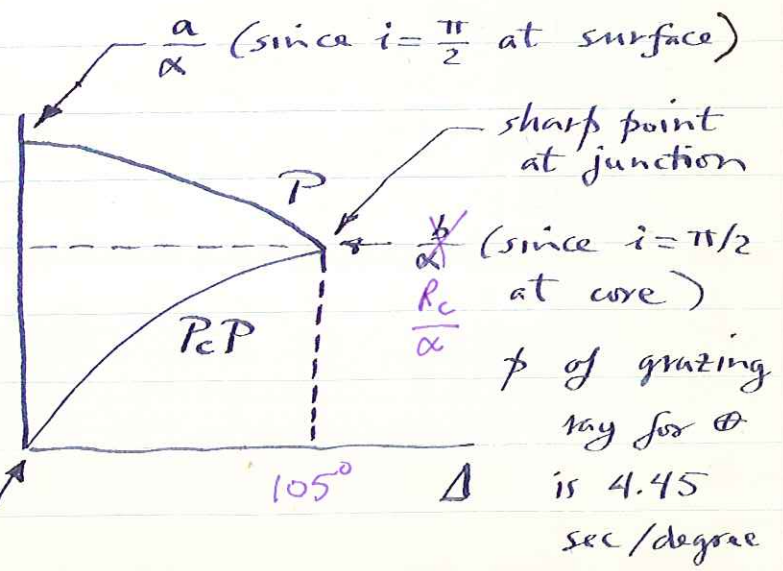
P and PcP merge into one grazing ray, edge of core shadow at $\Delta \sim 105^\circ$ for \oplus 's core.

emerges at $\Delta \lesssim 180^\circ$ for \oplus

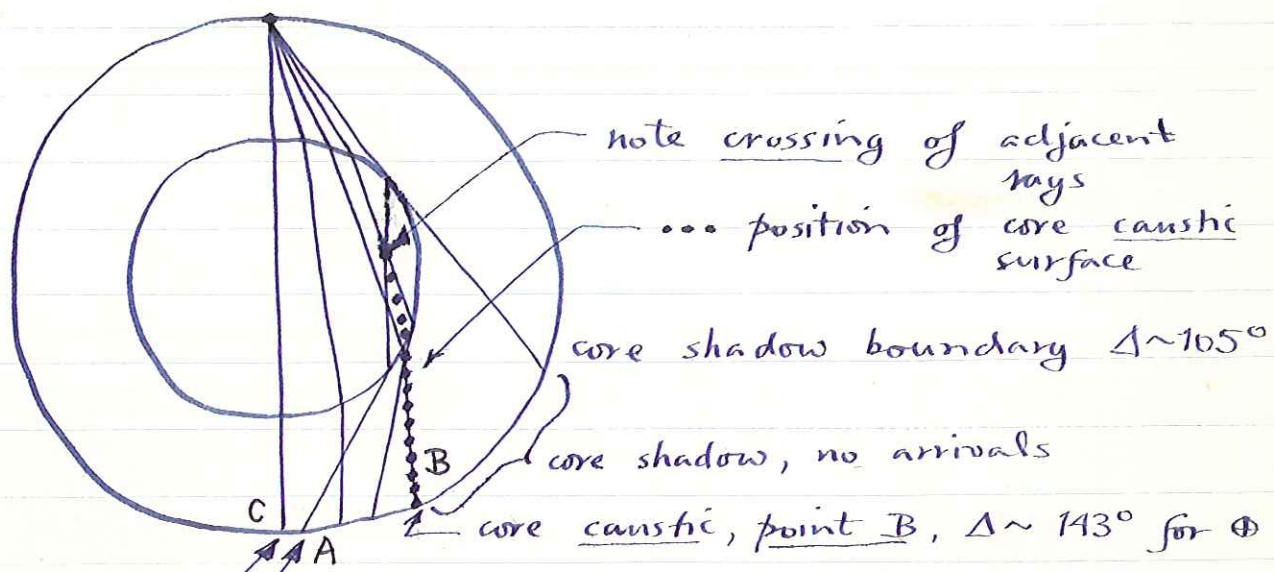


$\frac{dT}{d\Delta}$ decreases for P, increases from 0 for PcP

$\frac{dT}{d\Delta} = 0$, waves are coming straight up



For Δ beyond the core shadow: PKP waves refracted at core bdry.



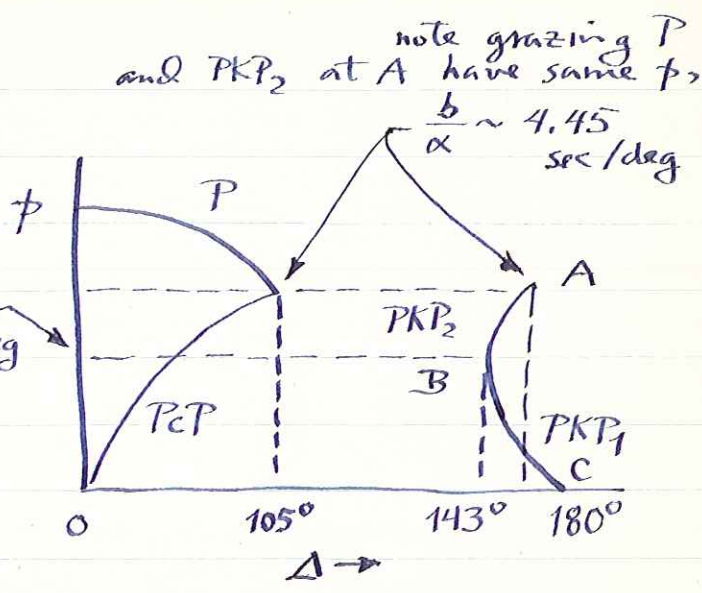
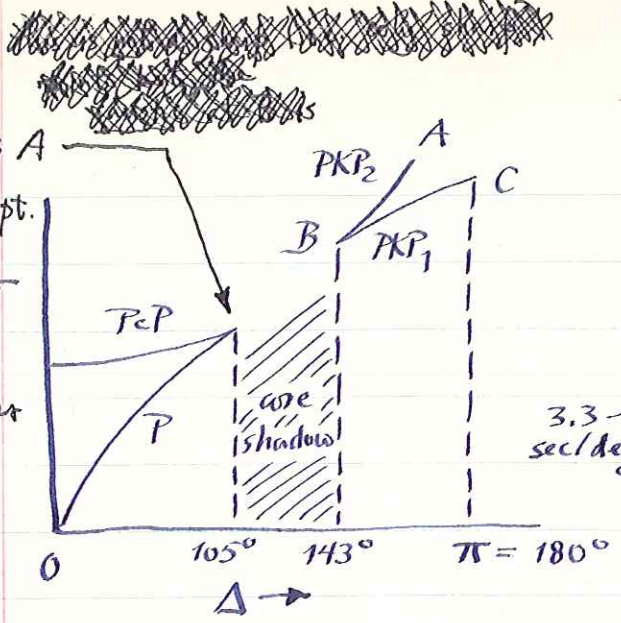
straight through $\Delta = 180^\circ$ point C, $p = \frac{dT}{d\Delta} = 0$

refracted ray at grazing incidence, point A, just less than $\Delta = 180^\circ$ for \oplus

The locus of crossing points of adjacent rays is called the core caustic surface. It can be visualized better in Fig. 9 from Julian + Anderson, shows P, PKP ray paths for a realistic \oplus model. I've dotted the crossing pts. of adjacent rays.

The $T(\Delta)$ and $p(\Delta)$ curves for the PKP waves have two branches in this simple model PKP_1 and PKP_2 .

this as well as is a sharp pt. since P and P_cP have same ray parameter ϕ .



Note: for PKP₂ both dT and dΔ are negative, so $\phi = dT/d\Delta$ is still positive.

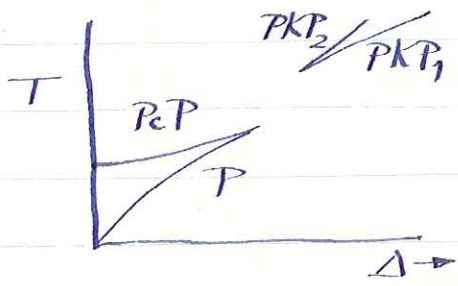
Caustics are recognizable by two features:

1. neighboring rays cross (actually become tangent, the caustic surface is the envelope of all possible rays)
2. $\phi(\Delta)$ turns vertical, not sharp point as at junction of P and P_cP.

Geometrical ray theory, in its simplest form as described here, fails to adequately describe signal at shadow bdy, the caustic point B and points A and C (the antipode) Diffraction effects become important in vicinity of all these pts, e.g. the shadow bdy is not sharp.

PKP in actual Earth:

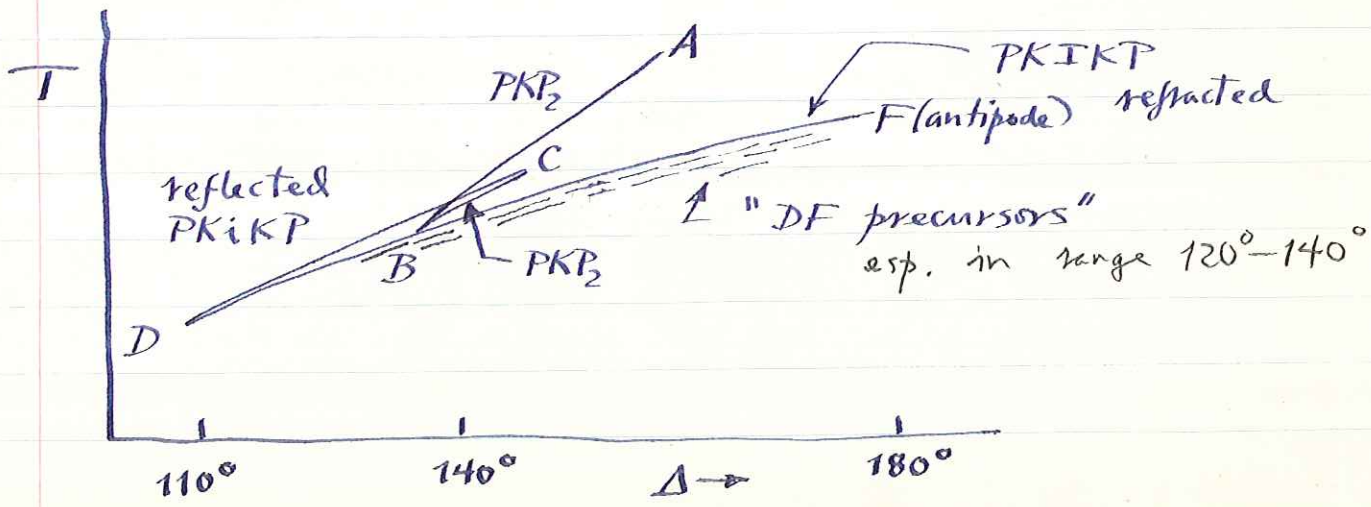
Oldham (1906): first evidence for low velocity core, observations of P, PcP, P_{diff} near the shadow. T(Δ) curve thus thought for many years to look like



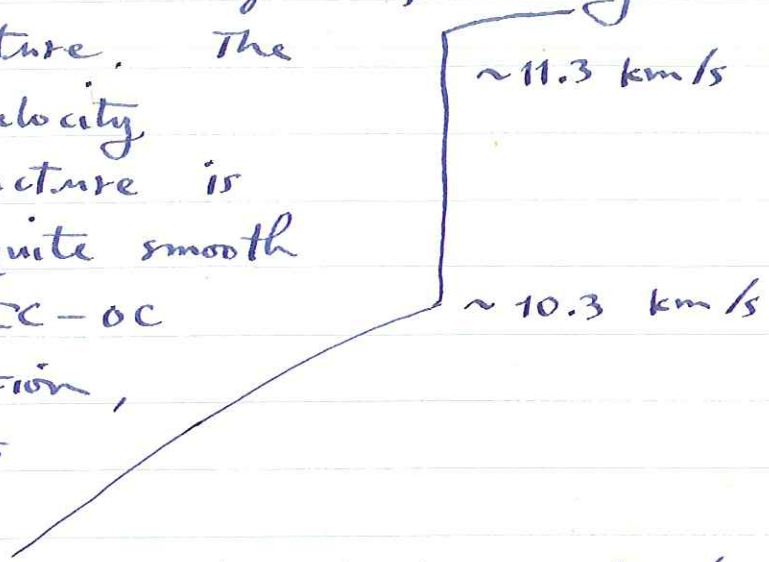
1936 Inge Lehman observed arrivals between 110° and 140° which could not be explained on this basis (they would be in the core shadow).

Interpreted as refractions from a high velocity inner core, also reflections from this bdry, Lehman the discoverer of the inner core.

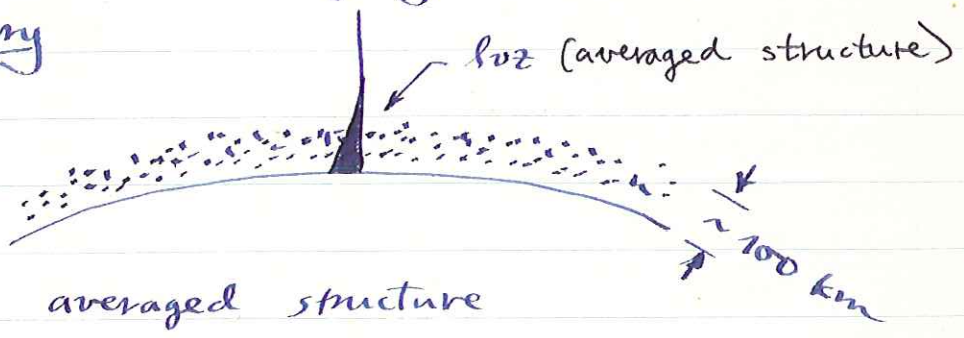
Travel time curve now known to look like:



core-mantle bdry. Above $T(\Delta)$ curve now thought to be quite reasonable for spherically averaged structure. The velocity structure is quite smooth at IC-OC transition, jumps from about 10 to about 11 km/s, much as envisaged by Ms. Lehman.

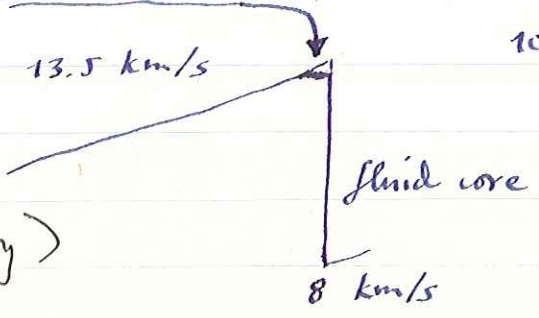


In contrast, core-mantle bdry, current picture: highly inhomogeneous above bdry

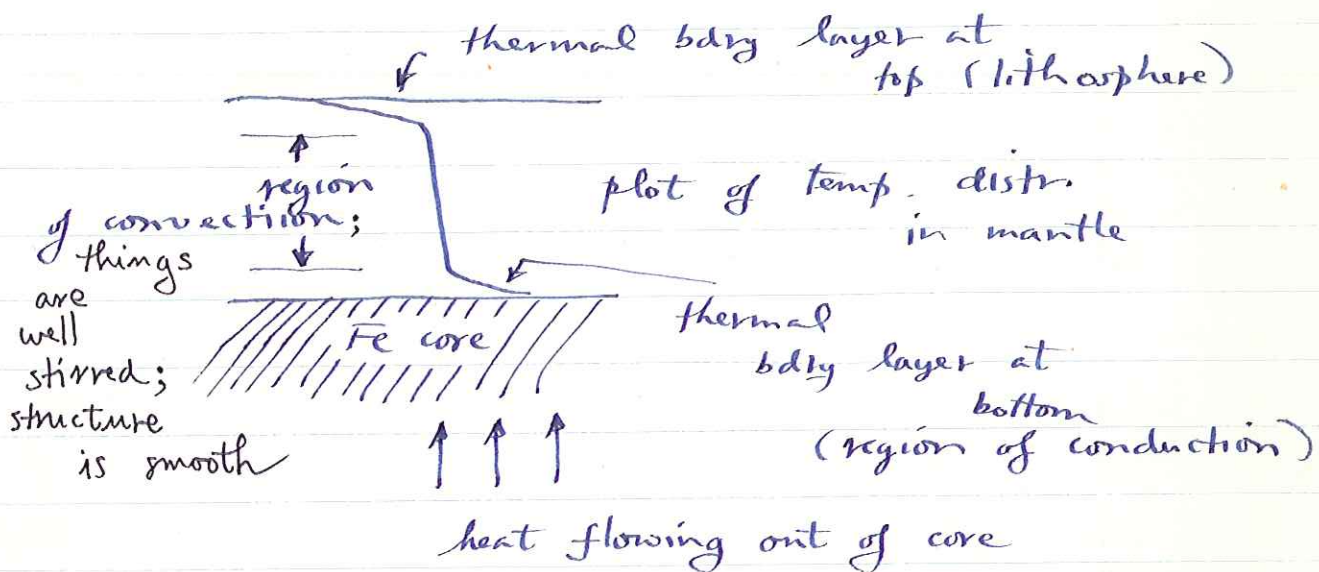


weak lvz, see e.g. model 1066A

Also suggestion of a Thorne layer (discontinuity)



Current interpretation, thermal bdry layer above core mantle bdry.



Underside of core-mantle bdry appears on other hand to be very smooth because travel times of P4KP, ~~P4KP~~ P7KP, P9KP, P11KP (!) etc. show no increasing scatter, arrivals are sharp and impulsive,

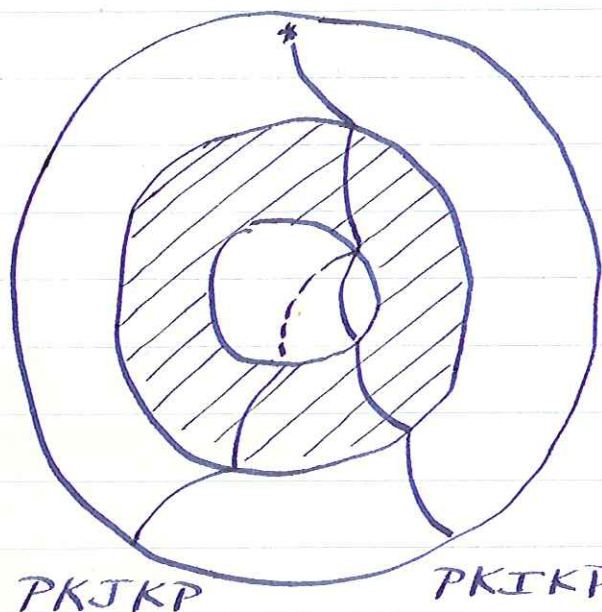
no min. marks
 → whole → length shown is less than a minute; this is a short-period record unlike those in our lab.

see example of P4KP cover of EOS. Any kumps on inside of core must be small ≤ 1 km, whether this large enough to perturb dynamo is debated.

PnKP rays really get thrown around, see e.g. P7KKP from Chapman, narrow solid angle at source gets sprayed out over $3/4$ \oplus surface ⇒ amplitude small.

Inner core long thought to be solid since this an easy way to $\uparrow \alpha(r)$.

Not until 1972 that a claim was made for observation of PKJKP (Julian, Sheppard, Davies) which would be direct proof of solidity.



Look for in
range

$$\Delta \sim 230^\circ - 290^\circ$$

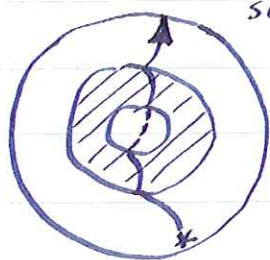
$$T \sim 1/2 \text{ hr.}$$

They found a
phase in \sim
right place.

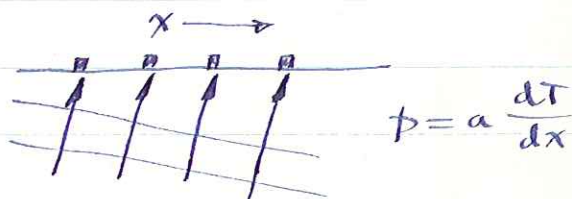
\exists a major unresolved problem with interpretation. They find $\beta \sim 3 \text{ km/s}$, cannot be reconciled with $\beta \sim 3.5 \text{ km/s}$ from inversion of free oscillations. Has been suggested they may have seen SKJKP, but they checked dT/dh using quakes of different depths. Its a fuzzle.

How are weak body phases looked for in general? Two tricks.

1. use deep focus quakes, not obscured by larger surface waves, we do this in our lab.
2. use beam steering of arrays



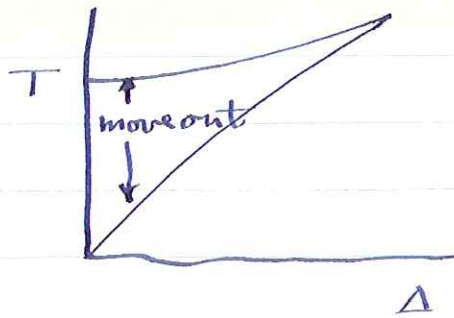
say want to look for a phase with a certain ray parameter p



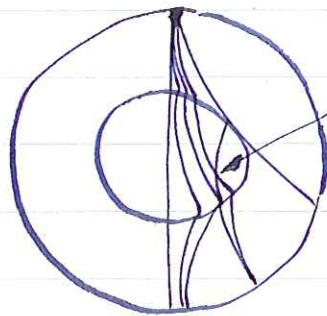
beam steering: simple delay and sum of signals at various seismometers.

Two additional comments:

1. PcP is generally quite a weak signal since reflection coefficient at CM bdy is small (~ 0.1) other 90% is transmitted to become PKP. See Fig. 12 from Buchbinder, disregard data, theoretical value from known contrast in properties, at near vertical incidence almost zero, we look for PcP in lab, can recognize by moveout at several stations, different Δ .



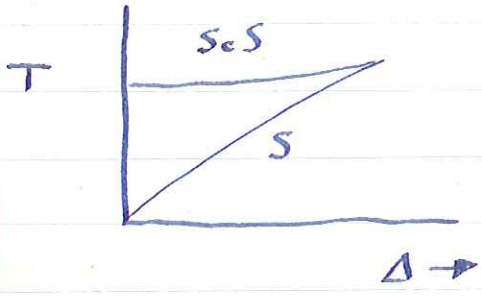
2. Note PKP does not bottom in upper part of core



no PKP's bottom in this region, they traverse it but don't bottom in it.

A ray is most sensitive to velocity near bottoming point. This makes it very difficult to use PKP times to find $\alpha(r)$ in outer half of core. We shall see that SKS does not have this problem, and can be (and is) so used. Instead it has a different problem: SKKKS, etc. all come in at \sim same time.

The conclusions regarding S rays in the mantle are similar, get S and ScS merging at core shadow, but (ScS)_{SH}



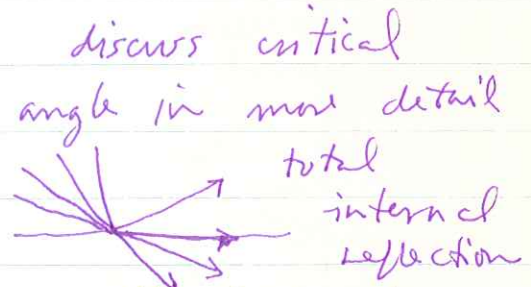
can be a strong arrival since its reflection coefficient is necessarily 1 (∴ no transmitted SH)

The refracted core waves have quite a different behavior since

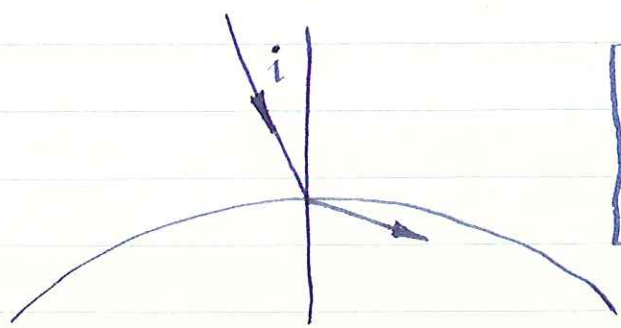
$$\beta_{\text{mantle}} \sim 7.2 \text{ km/s}$$

$$\alpha_{\text{core}} \sim 8.0 \text{ km/s}$$

$$\beta_{\text{mantle}} < \alpha_{\text{core}}$$



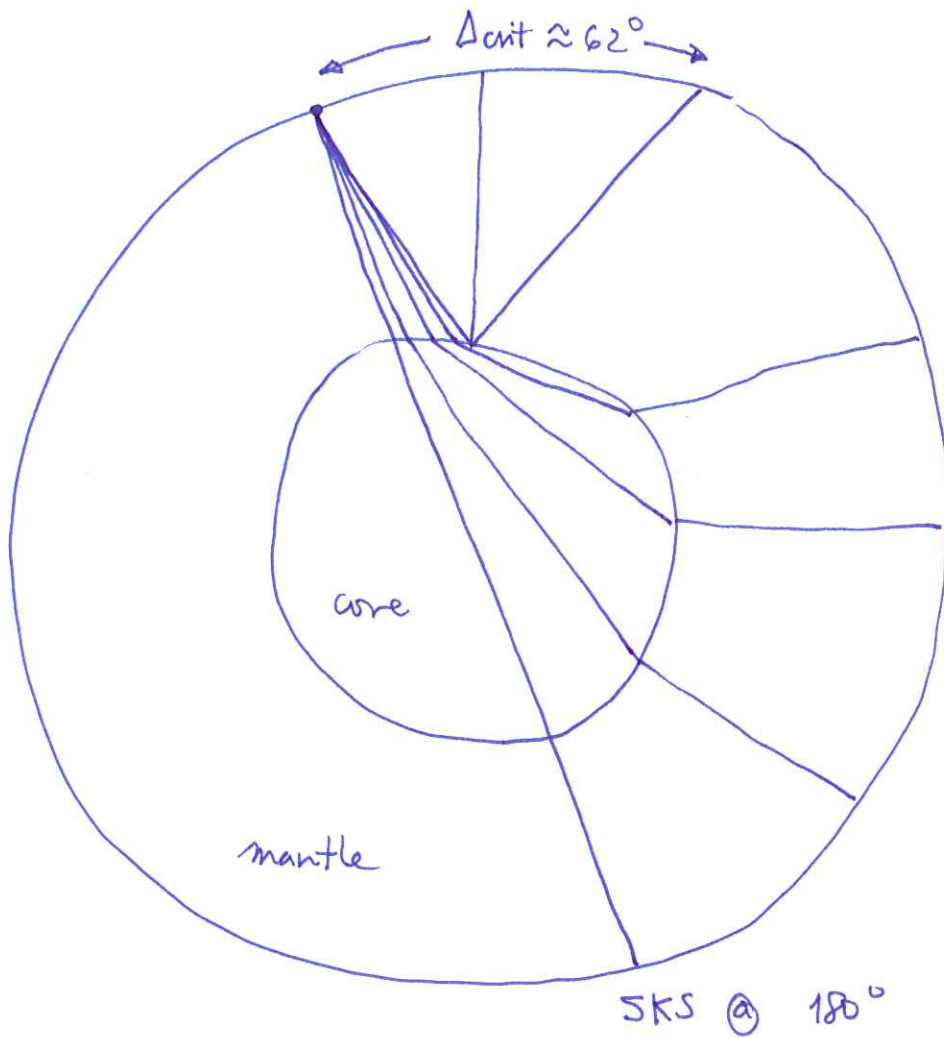
PKP core rays existed for all incident angles, but now there is a critical angle i_c beyond which no rays are refracted into the core.



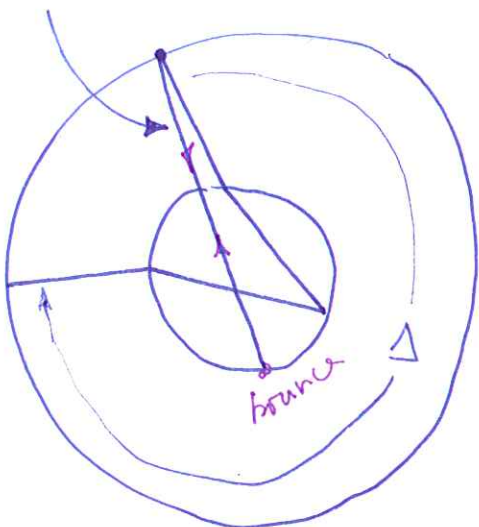
$$i_{\text{refr}} = 90^\circ$$

$$\sin i_c = \beta_{\text{mantle}} / \alpha_{\text{core}}$$

Consider possible SKS rays, starting from 180° :

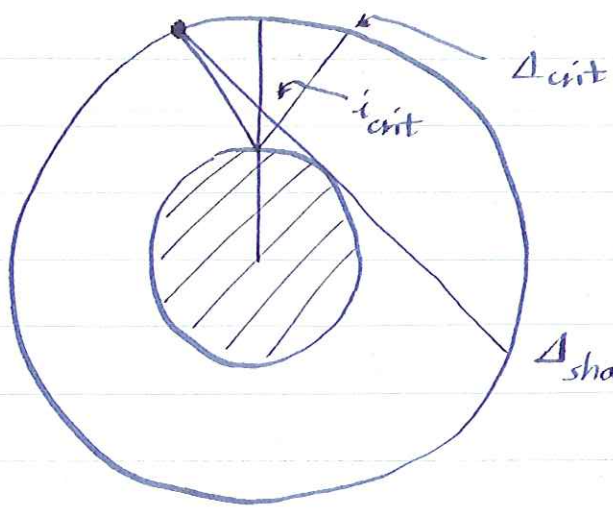


SKKS ~~SKKS~~ @ $\Delta = 360^\circ$



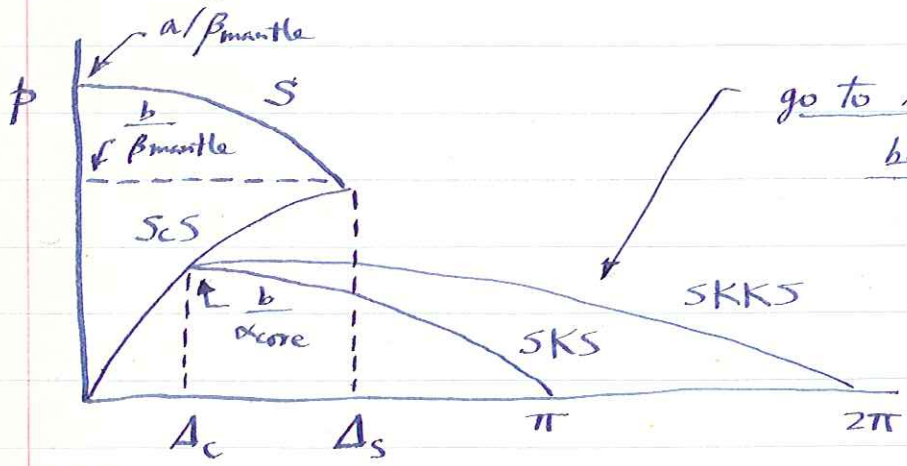
SKKS
~~SKKS~~ at angles $> 180^\circ$ look like this

Thus we have, for the homogeneous core and mantle case considered before,



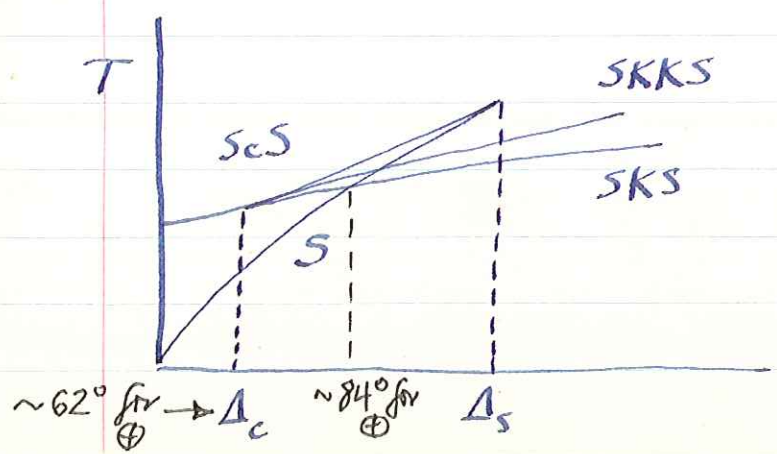
No SKS for $\Delta < \Delta_{crit}$
 $\Delta_{crit} \sim \text{old}$ for \oplus
 62°

For $\Delta = \Delta_{crit}$, SKS and ScS coincide. Plot of \dot{p} vs. Δ thus looks like:



go to next page before drawing SKKS, etc. in this figure and that below

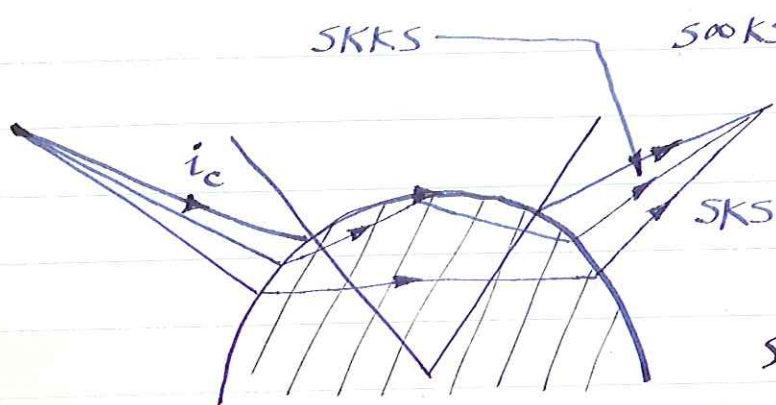
The travel time curve looks like:



$\Delta \sim 84^\circ$ where S and SKS cross.

What about SKKS, SKKKS, etc.?

At any $\Delta > \Delta_{crit}$ \exists an ∞ family of rays SKS, SKKS, etc.



S00KS suffers an ∞ no. of internal bounces, travels along core-mantle interface

Similar to the whispering gallery phenomenon in St. Paul's Cathedral, London.

The times of the rays SKKS, SKKKS, etc. approach that of S00KS which appears to travel along the core-mantle interface

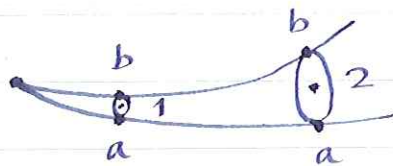
Geometrical ray theory clearly fails adequately to describe this coalescence of rays.

Ray tracing plot for SKKKS: note small azimuth range leaving source gets spread out considerably, amplitudes of SKKKS etc. thus weak.

S00KS has $n-1$ caustics

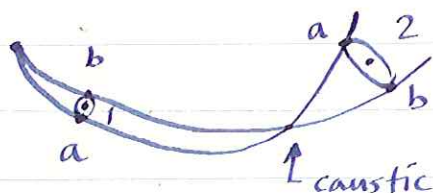
Note caustic surface inside core: locus of crossing points of nearby rays, this surface does not intersect surface of \oplus .

Passage of a ray through a caustic gives rise to a $\pi/2$ change in phase.
Reason: consider ray bundle



Amplitude law:
 $A_2/A_1 \sim (\text{area 1}/\text{area 2})^{1/2}$

Caustic surface is where nearby rays cross



The ray bundle gets turned inside out. The area at 2 is negative.

Thus (this actually just a mnemonic device, not a rigorous derivation):

$$A_2/A_1 = \left(- \left| \frac{\text{area 1}}{\text{area 2}} \right| \right)^{1/2}$$

$$= i \left| \frac{\text{area 1}}{\text{area 2}} \right|^{1/2}$$

, $i = e^{i\pi/2}$
a $\pi/2$ phase advance upon passage thru caustic

This not easy to observe with SKKS, SKKKS, etc. since \exists phase shifts at CM bdy as well which depend upon angle i .

But another such situation of an internal caustic is PP, SS etc.

See ray tracing plot for PP; SS would look similar. If attention restricted to SH phase shift upon surface reflection is nil and SS should be phase advanced relative to S by exactly $\pi/2$.

Given a pulse $s(t)$ if every constituent Fourier component is phase shifted uniformly by $\pi/2$ the result is said to be the Hilbert transform of $s(t)$. We may think of passage thru a caustic as a black box which converts $s(t)$ into $H[s(t)]$, its Hilbert transform.

$$\sin \omega t \rightarrow \boxed{\text{Hilbert transform}} \rightarrow \cos \omega t = \sin(\omega t + \frac{\pi}{2})$$

Passage thru 2 such boxes advances phase by π , equivalent to multiplication by -1

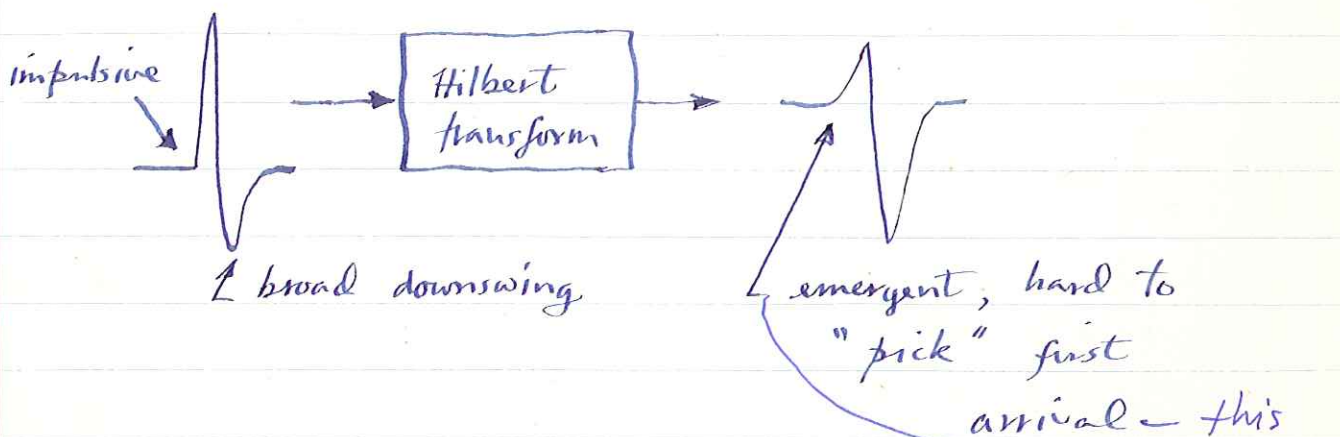
$$s(t) \rightarrow \boxed{\text{Hilbert transform}} \rightarrow \boxed{\text{Hilbert transform}} \rightarrow -s(t) = \sin(\omega t + 180^\circ)$$

in the case above

Figs. 9 and 10 from Choy + Richards show

two transverse $S + SS$ seismograms.

The Hilbert transform of a seismograph-filtered impulse looks like the signal shown, viz.

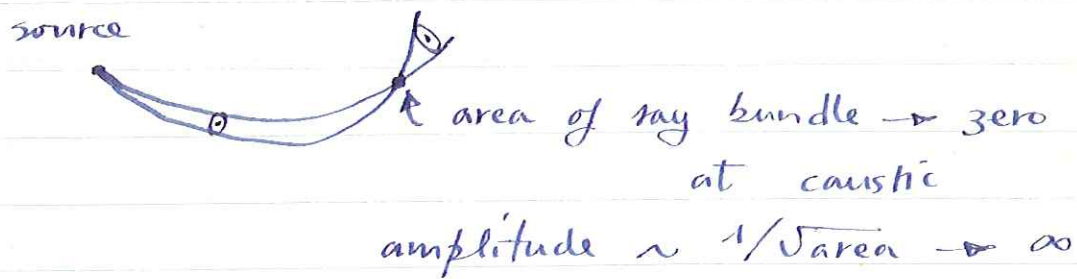


But ~~the~~ Hilbert transformation of the seismogram, accomplished easily in computer, has precisely the expected effect:

1. $H[S]$ takes on the emergent character of SS
2. $H[SS]$ looks like $-S$, but weaker due to shear attenuation, finite $Q\beta$.

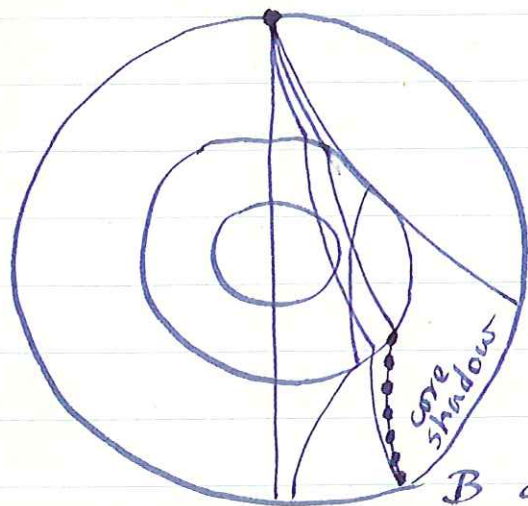
This a particularly clear example of caustic phase shift.

Note that at the caustic itself ray theory predicts an ∞ amplitude



This an example of failure of ray theory, fails near caustics, focal points and shadow boundaries.

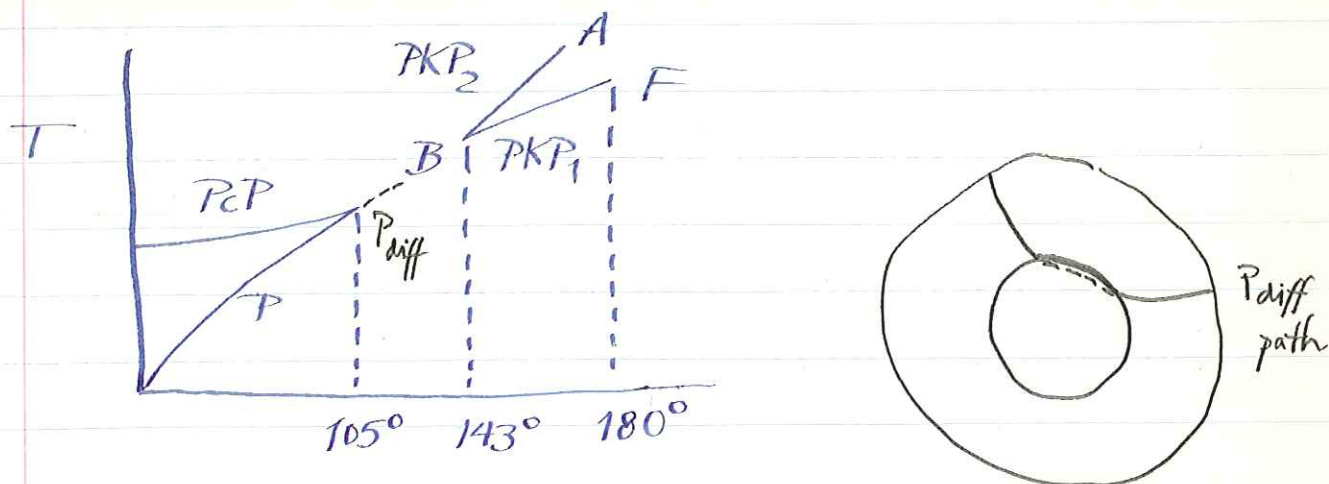
At all such locations the finite frequency of the waves must be taken into account. Above problem does not arise for SS etc. since caustic does not intersect \oplus surface, but PKP caustic does, at $\Delta \sim 143^\circ$.



A core shadow $\Delta \sim 105^\circ$

B caustic $\Delta \sim 143^\circ$

F antipode $\Delta = 180^\circ$

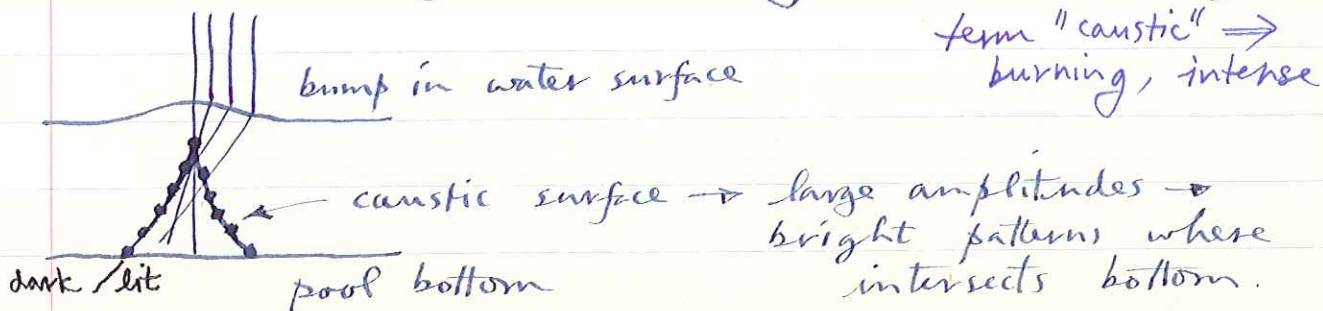


Ray theory predicts ∞ amplitude arrivals at point B, PKP caustic, also predicts sharp shadow bdy at $\Delta \sim 105^\circ$.

Actually diffracted P_{diff} is seen far into the shadow, particularly at long periods.

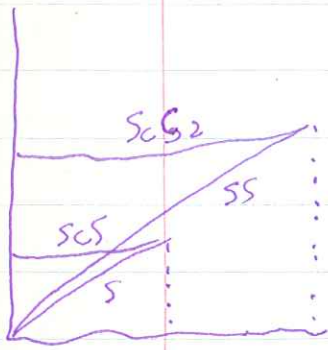
Amplitude at B is large but not ∞ .
Theory for behavior of waves in vicinity of caustic due to Airy, whom we have met in connection with isostasy.

An everyday example of caustics: patterns of dark and light on bottom of a swimming pool on sunny day, due to refraction at air-water interface, focusing and defocusing



Travel time plot for surface focus shows the various phenomena associated with the core we have been discussing :

1. P and P_{diff} joining at core shadow, P_{diff} extending way into core shadow, makes it difficult to determine location of shadow boundary

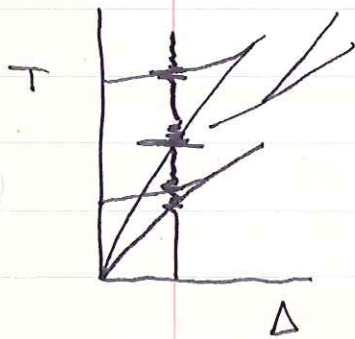


2. PKP₁ and PKP₂ with CD branch (PKiKP) going into core shadow
3. notice also PP, PPP, etc.
4. S and ScS, SKS, SKKS, etc.

All the above are empirical. Many of the phases shown not necessarily observed. Easy to determine expected times of compound phases e.g. PKPPKP by adding relevant segments together

5. LQ and LR, straight lines, constant velocity are approximate arrival times of 20 s period Love and Rayleigh surface waves, dominant signal for oceanic paths, shallow focus events.

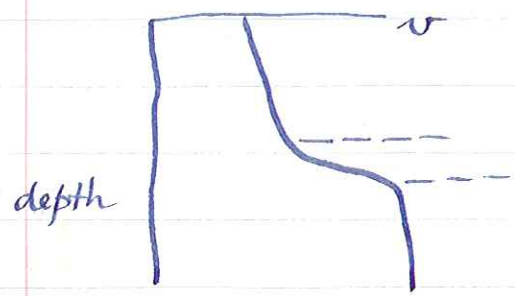
expect many arrivals



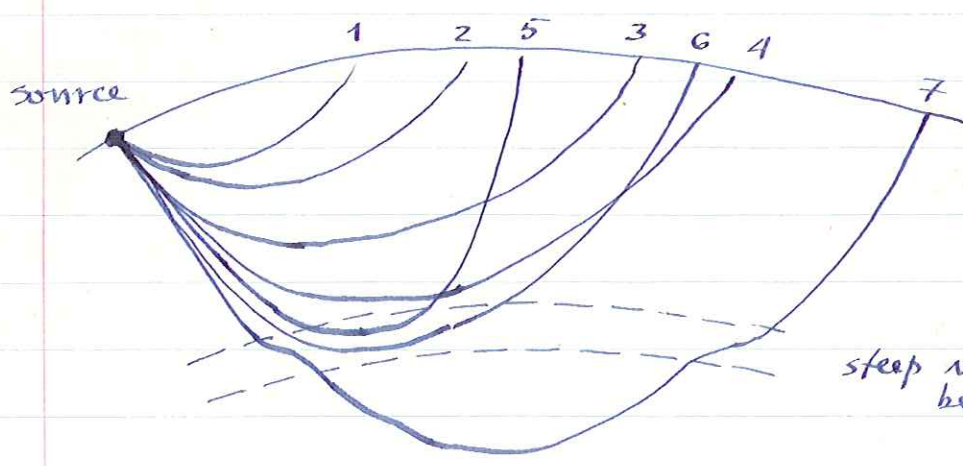
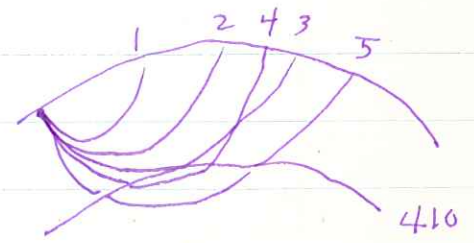
Consider now effect of continuous but rapid changes in velocity.

Also important in \oplus , two cases of ~~interest~~ interest, both occur in upper mantle.

High velocity gradient: similar to S, SKS case. Suppose $v(r)$ looks like

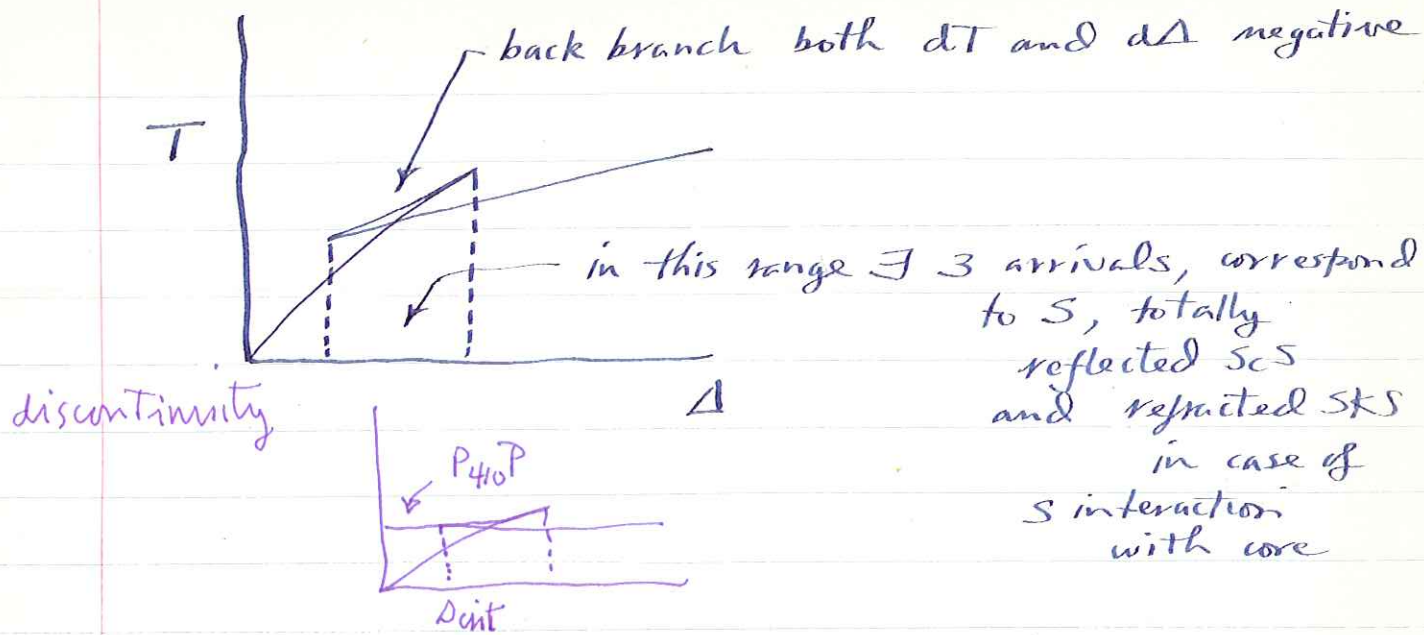


steep gradient region
upper-mantle discontinuity

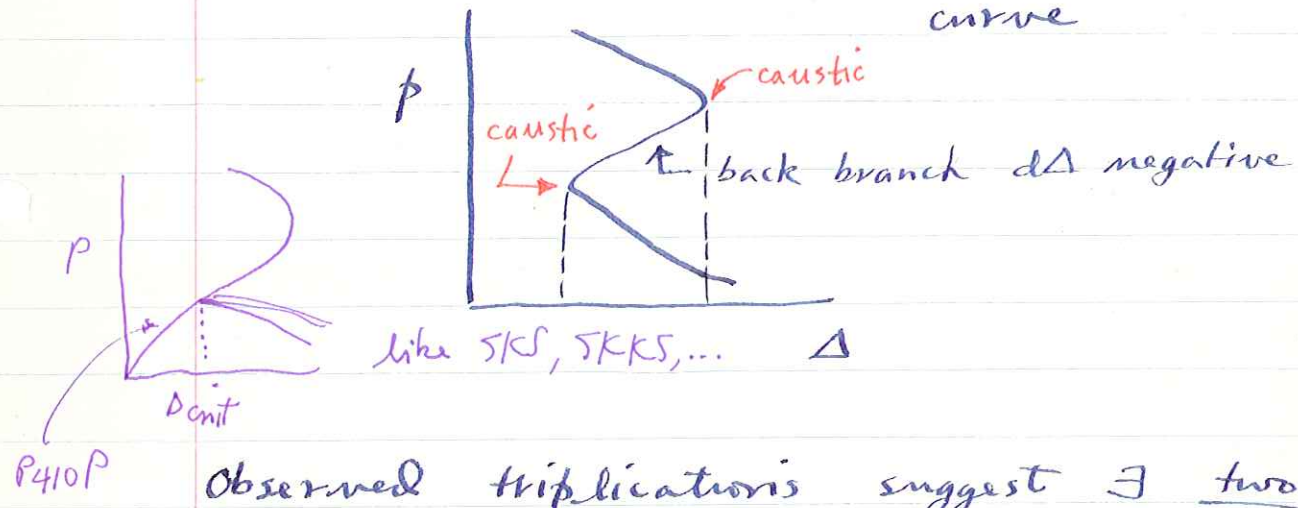


steep velocity gradient bends rays sharply

Travel time plot thus looks like:
called a triplication.



$p(\Delta)$ thus looks like, smooth continuous curve



Observed triplications suggest \exists two major high velocity gradient zones in upper mantle at depths of ~ 410 km and ~ 650 km. Easiest to observe using arrays. Recall Fig. 4 from TFSO study by Lane Johnson, location of epicenters used shown in Fig. 5

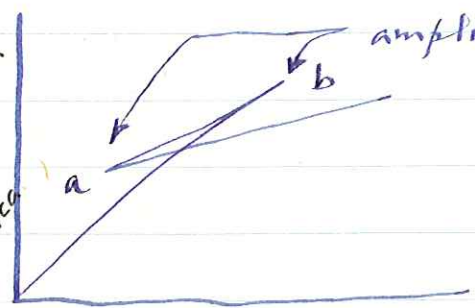
One triplication in range $15^\circ \lesssim \Delta \lesssim 20^\circ$ and another in range $20^\circ \lesssim \Delta \lesssim 25^\circ$.

His interpretation CIT204 shown in Fig. 24, Two steep gradient zones at ~~xxxx~~ ~ 400 km and ~ 650 km depth.

To actually see all three arrivals is not easy. Look at ray tracing diagram for Δ out to 40° for model CIT204 from Julian and Anderson Fig. 25.

Spacing of rays at surface an indication of expected amplitude (if shot out from source in equal solid angles). In general for a triplication it is found that

Look also at
Walck &
Clayton section - second
Army - Middle America
recond - Scarlet
quakes



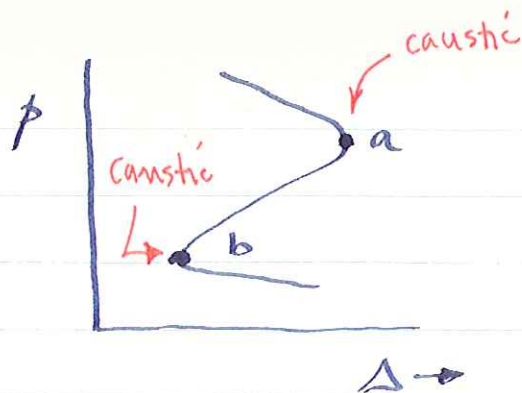
amplitudes
largest at point a
(corresponds essentially
to total internal
reflection)

next largest at
b, weaker between
a and b

Both a and b are
caustics (for a smooth $v(r)$); $f(\Delta)$ is smoothly

In fact this another example of failure of vertical ray theory, which predicts that

$$\text{amplitude} \sim (d^2T/d\Delta^2)^{1/2}$$



Ray theory predicts ∞ amplitudes at points a and b since

$$d^2T/d\Delta^2 = dp/d\Delta$$

$\rightarrow \infty$ (p vs. Δ turns vertical)

For a model get, instead, $v(r)$ change in slope

Modern Δ method of studying upper mantle structure is using synthetic seismograms, try to match amplitudes of later arrivals and other details of waveform.

Example of this method: work of Helmberger and Wiggins LRSM stations in western U.S., nuclear explosions at NTS,

Fig. 2 shows two record sections reduced to velocity of $10.8 \text{ } \frac{\text{degrees}}{\text{second}}$.

Large amplitude second arrival in range $\Delta \sim 15^\circ - 20^\circ$ very clear, see e.g. station WNSD from Aardvaark, due to "discontinuity" at 400 km depth.

Fig. 16 shows their suggested refinement of CIT204 (---), p vs. Δ curve with more structure