

# Picking first arrivals in hydroacoustic seismograms from MERMAID floats

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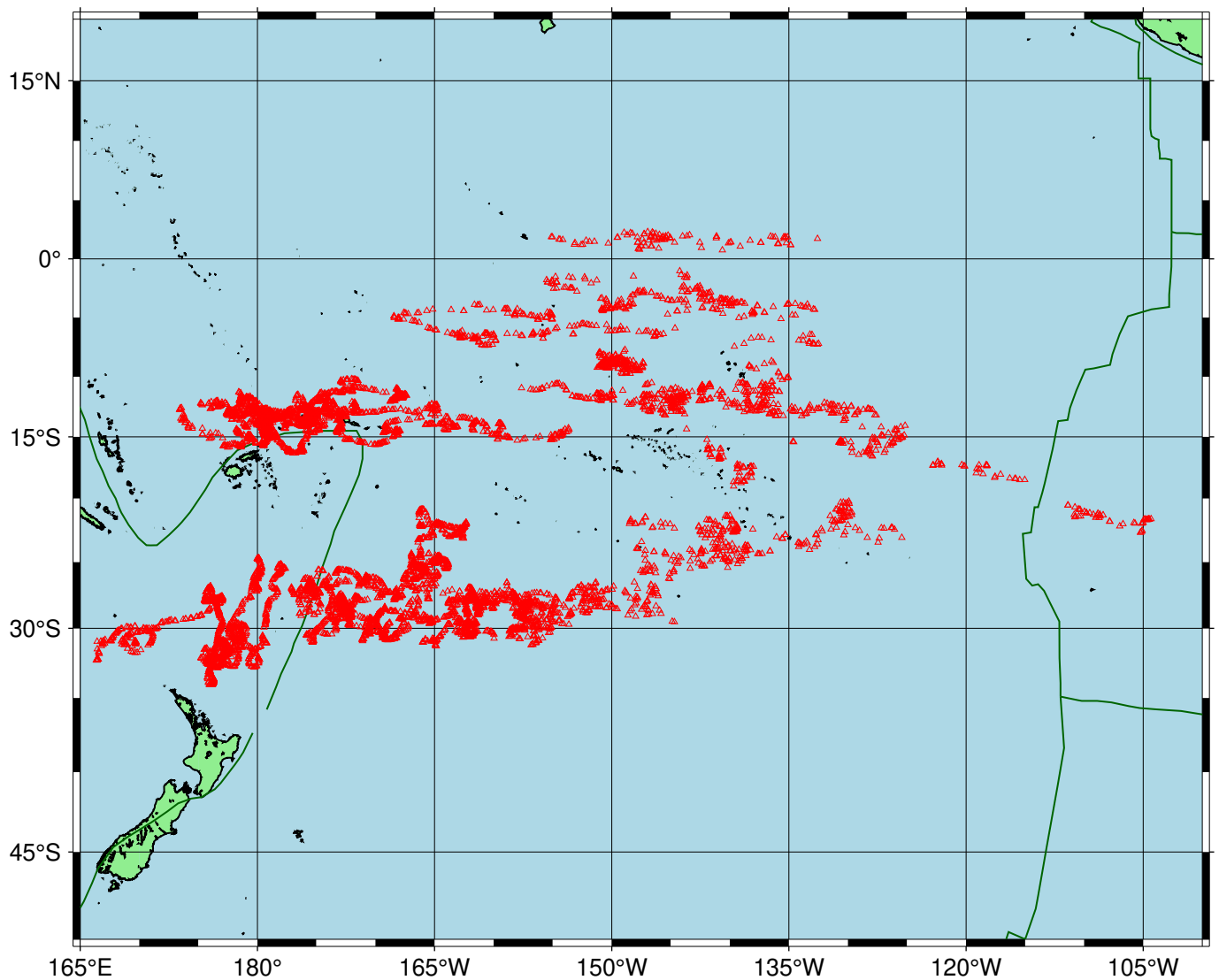
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**Abstract** Floating seismometers ('MERMAIDS') operating in the noisy environment of the world's oceans pose a challenge for picking the time of earthquake first arrivals. We report on an experiment to estimate the errors in picked arrivals from 49 MERMAIDS operating in the South Pacific, using two independent strategies. For 15 events, the same arrivals were redundantly picked by several analysts, allowing for a direct estimate of error distributions. Standard errors in times from MERMAID seismograms vary from 0.2 s for close events at mantle depths in the Kermadec subduction zone to more than 2 s for crustal events at large epicentral distance. In a second experiment we analysed the a posteriori misfits after tomographically inverting all events. The residual traveltime misfit is consistent with the error estimates from the first experiment, but also shows inconsistencies with arrival times from the ISC-EHB and NEIC catalogues, which we attribute to errors in the published hypocentres and/or origin times.

## 1 Introduction

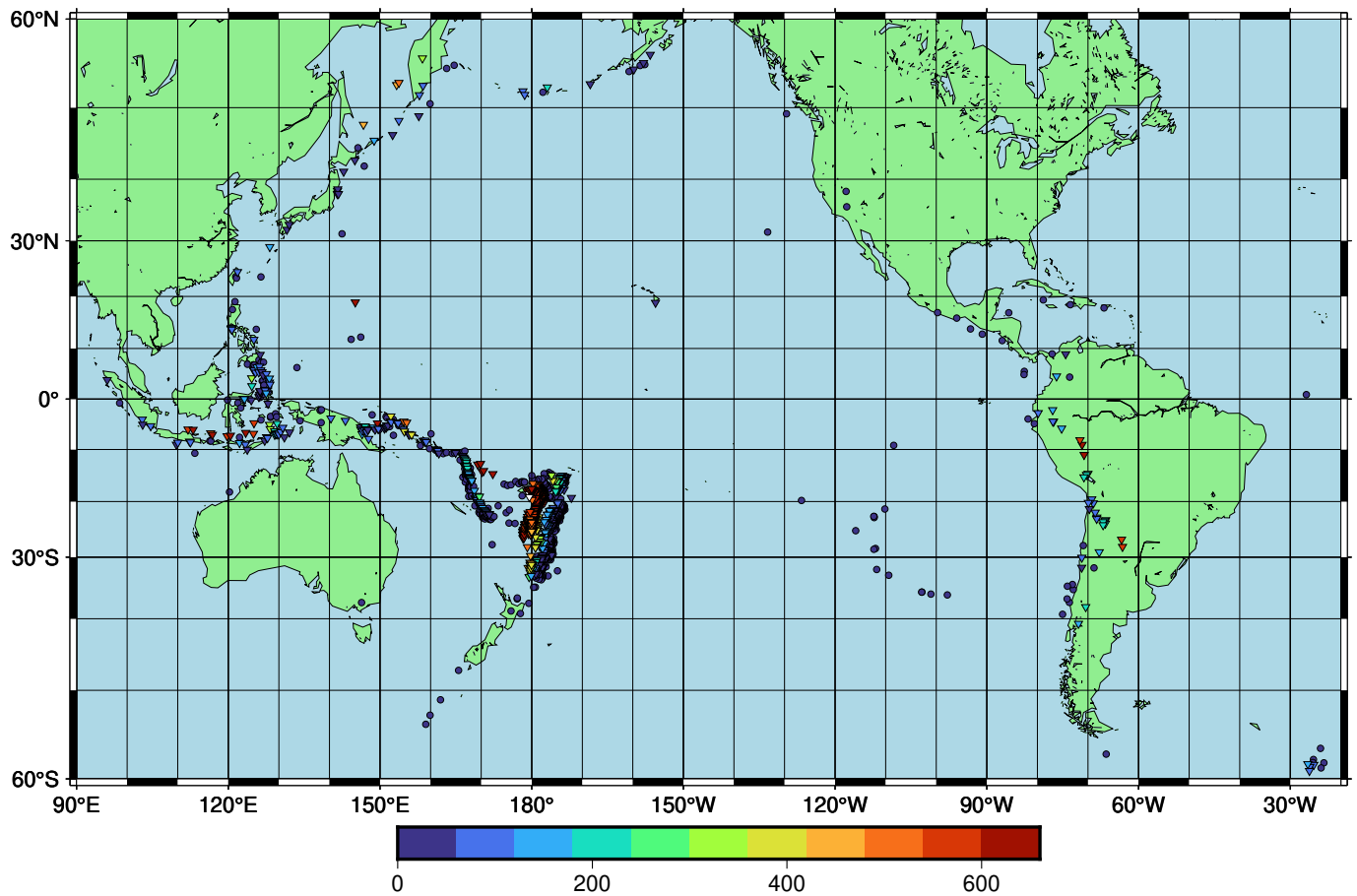
MERMAIDS or 'Mobile Earthquake Recording in Marine Areas by Independent Divers' (Simons et al., 2009) drift passively deep below the ocean's surface (typically at 1500 m) and are equipped with a continuously recording hydrophone. The pass band of the instrument is between about 0.05 Hz to the Nyquist frequency of 10 Hz, though only local events generate significant signal above 2 Hz. A triggering algorithm (Sukhovich et al., 2011) keeps track of presumed P-wave arrivals. For sufficiently strong signals, it commands the float to rise to the surface, transmit the most recent recording with a latency of several hours, depending on the rise time, together with possible weaker P-arrivals stored earlier. The location where the actual recording took place is determined by interpolation of GPS fixes (Nolet et al., 2024).



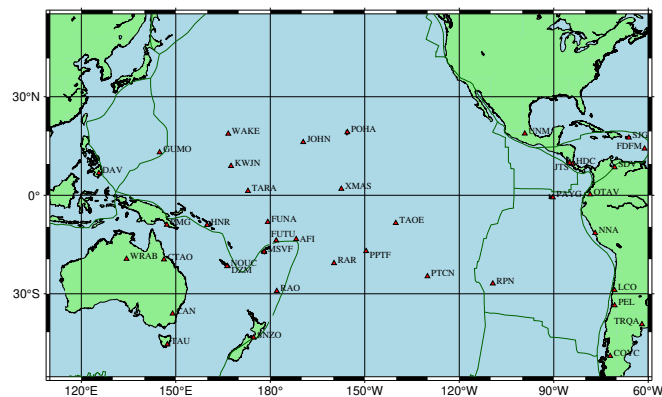
**Figure 1** Locations where MERMAIDs recorded the seismograms used in this study. Note the increased density in the western part of the domain, where many very weak Tonga-Fiji and Kermadec events occurred close enough to the instruments to have an acceptable signal-to-noise ratio. Plate boundaries are indicated by thin black lines.

33 The noise level in the seismograms is generally high, rendering the picking of first arrivals difficult. We have  
34 developed a first-arrival picking strategy based on comparison with other MERMAIDs and nearby island stations,  
35 knowledge of the expected polarity of the P-wave, and using both broad-band and high-pass filtered records. The  
36 data processing of the MERMAID seismograms includes an initial arrival time estimate using the Akaike Informa-  
37 tion Criterion, or AIC (Simon et al., 2020) which is often – but not always – within about 0.2 s of the visual pick by  
38 the analyst. Since MERMAIDs are relatively recent additions to the seismological toolbox, not enough data are yet  
39 available to train an AI-based algorithm (Mousavi et al., 2019; Lomax et al., 2024), though we hope that the current  
40 effort will take us many steps in that direction.

41 The data in this study are from 49 MERMAIDs in the South Pacific, of which the earliest were launched in June  
42 2018 as part of the SPPIM, or ‘South Pacific Plume Imaging and Modeling’, project (Simon et al., 2020, 2022). The  
43 large majority of floats is still operational today and in this paper we use data transmitted until November 2023. A  
44 live map of the state of the network is available on the web ([www.earthscopeoceans.org](http://www.earthscopeoceans.org)), where one can also inspect  
45 the history of each float. All data are being archived by the EarthScope Consortium with network code MH.



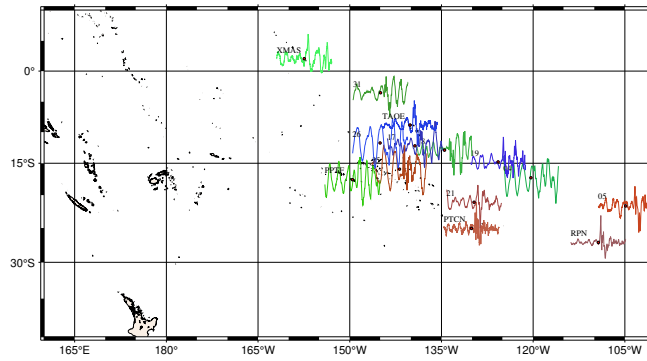
**Figure 2** Locations of 1147 deep events (triangles) and 703 shallow ones ( $h < 35\text{km}$ , circles) analysed in Experiment 2, with colour indicating the hypocentre depth in km.



**Figure 3** Locations of the 40 surface (land) stations used to compare waveforms.

46 Our team of ‘pickers’ consists of experienced seismologists, PhD students and postdocs from institutions partic-  
 47 ipating in SPPIM (see author list). Prior to doing the experiment, a series of training sessions was held via Zoom, of  
 48 which the materials are available on the web for future users of MERMAIDs (Nolet, 2024). Since the ultimate goal of  
 49 SPPIM is to sharpen tomographic images of the upwelling mantle structure(s) beneath the South Pacific, a correct  
 50 estimate of picking errors is essential.

51 Ideally, we would like to see picking errors well below the uncertainty introduced by the crustal corrections in  
 52 tomography and possible errors in the location of the float. In our case the error in crustal corrections is dominated by  
 53 the uncertainty in satellite bathymetry, which cannot account for rapid changes at wavelengths  $< 10\text{ km}$ . Sepúlveda.  
 54 et al. (2020) give an estimated standard error of 160 m for satellite bathymetry near Chile. Our own knowledge of



**Figure 4** All MERMAID seismograms for the M 6.5 event in the Peru-Brazil border region of 2022/06/08 plotted on a map, together with nearby surface stations. The red dot indicates the station location, and is also the expected time of arrival of the P wave for model AK135.

55 bathymetry errors is largely anecdotal, but it has not been unusual to see a MERMAID float happily at 1500 m where  
 56 the bathymetry from GEBCO 2014 (Weatherall et al., 2015) reports less than a one kilometre of water depth, indicating  
 57 the error may be significantly larger than 160 m in the SPPIM area. We must add to this the error in the corrections  
 58 for the Moho depth and oceanic crustal structure, for which we used the crustal part of LITH1.0 (Pasyanos et al., 2014).  
 59 We therefore assume a prior uncertainty in the total crustal correction of 0.4 s in tomographic inversions.

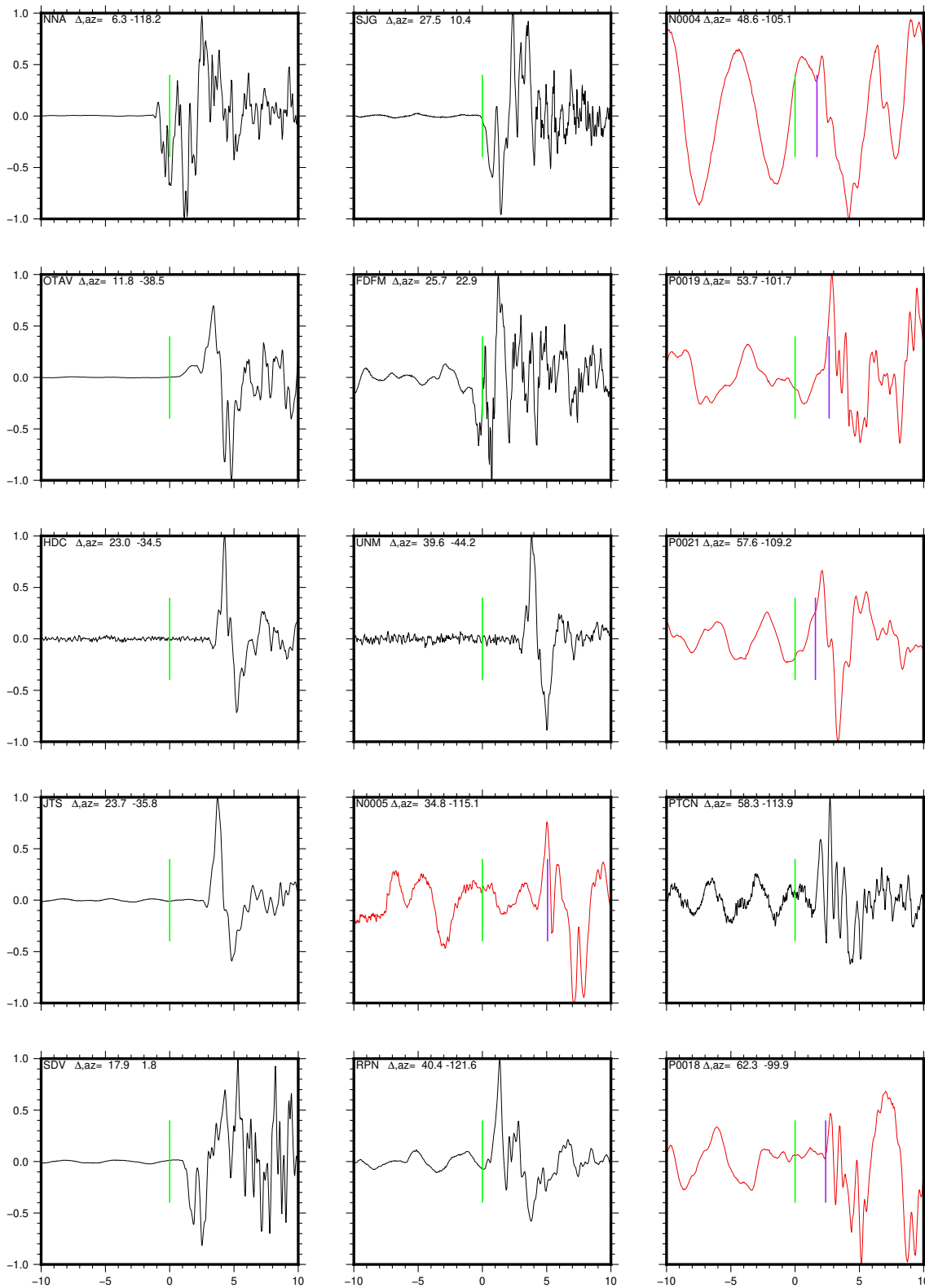
60 The equivalent travel time delay error caused by mislocation of the float is generally below 0.1 s (Nolet et al., 2024).  
 61 Contrary to OBS data, we do not have to worry about clock correction errors (Naranjo et al., 2024) since the internal  
 62 clock drift is regularly measured, and corrected for, by GPS at each surfacing.

63 To estimate the picking errors we develop two strategies. In the first experiment, seismograms from MERMAIDS  
 64 as well as nearby island stations are picked multiple times by different pickers. In the second experiment, we invert  
 65 the 16,739 arrival time picks from 1850 events and measure the a posteriori fit to the predicted times.

## 66 2 Data

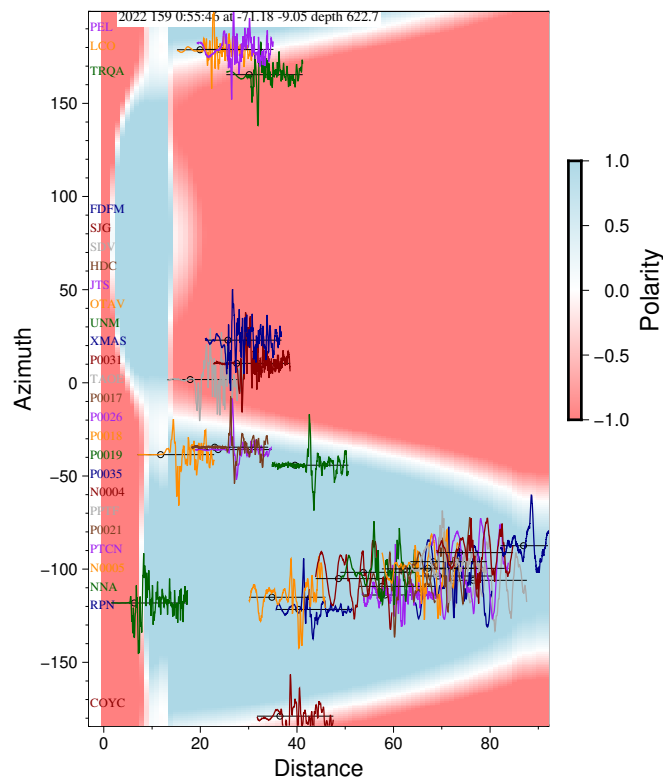
67 The 49 MERMAIDS in the SPPIM project were launched from scientific vessels operated by Ifremer in France and  
 68 JAMSTEC in Japan. The first float (P0006) was launched on June 26, 2018. Two more cruises followed until the network  
 69 was complete by September 2019. At the time of writing this paper in the fall of 2024 all floats have thus exceeded  
 70 their designed battery lifetime of 5 years, and 36 of them are still operating, including P0006, which attests to the  
 71 durability of the instrument. Figure 1 shows the locations where MERMAIDS recorded a seismogram from one of  
 72 the 403 earthquakes shown in Figure 2. We used 40 surface (or borehole) stations from the global seismic network  
 73 to compare waveforms (Figure 3). Arrivals at these stations are picked as well, such that we are able to compare  
 74 the quality of picks from surface stations with those from the MERMAIDS. In total, we assembled 5384 picks from  
 75 MERMAIDS and 11,355 from surface stations. The addition of land station picks is also done to be able to apply event  
 76 relocations and origin time corrections at the time of inversions, since the number of MERMAID picks can be very  
 77 limited for low magnitude events only recorded by nearby floats. For all events, hypocentre metadata are taken from  
 78 the ISC-EHB catalogue when available (i.e. until 2020). For more recent events we use the latest NEIC estimates.

79 We have developed a highly streamlined procedure to pick first arrivals, implemented as Linux shell scripts. Fig-  
 80 ures 4 – 6 show the diagnostic screen output an analyst is presented with prior to picking an event. All seismograms



**Figure 5** Seismograms for the event of 2022/06/08 plotted in an order that allows for easy comparison of waveforms in nearby stations. Epicentral distance  $\Delta$  and azimuth are plotted in the upper left corner. The green line indicates the expected P wave arrival (using model AK135), purple lines those of the AIC pick from ? in MERMAID seismograms. To distinguish them from surface stations, MERMAID seismograms are coloured red.

81 for the event are plotted on a map to enable visual comparison in geographical context (Figure 4), as well as com-  
 82 bined in one plot in an order that allows for easy comparison of nearby stations (Figure 5). The most useful plot is  
 83 that of the predicted polarity (Figure 6), using published moment tensor estimates. Whenever available, we use the  
 84 SCARDEC double-couple tensor (Vallée et al., 2011), since it is more representative for the high-frequency arrivals



**Figure 6** Predicted polarity (UP=blue) for event 2022/06/08 (71°W, 9°S, 622 km depth). Small circles indicate distance and azimuth, plotted at the expected arrival time of the P wave.

85 that we target than centroid estimates (Ekström et al., 2012; Rösler et al., 2023, 2024). In any case, we inspect the  
 86 plot for any systematic deviations from the predicted polarity – which occur especially near the (white) nodal zones.  
 87 Unless the prediction is ambiguous, we only pick an arrival that has the predicted polarity as read from this plot.  
 88 Finally, we also inspect SCARDEC source time functions (Vallée and Douet, 2016), whenever available.

89 Once this initial orientation complete, the Seismic Analysis Code (SAC) program (Goldstein et al., 2003; Goldstein  
 90 and Snoke, 2005) is called up and seismograms are shown one after the other in a sequence that tries to optimize  
 91 nearby seismograms to follow each other. Figure 7 shows an example of such plot, offering the seismogram both as  
 92 a record high-passed at 1 Hz (using a one-pass Butterworth filter with only two poles, which produces a rather gentle  
 93 damping of lower frequency), and as the original broadband record. To help identify the P-arrival in the presence of  
 94 noise, the arrival time predicted by AK135 ("P") and the AIC estimate of the arrival ("F") are superimposed as vertical  
 95 lines. The latter detects where the variance of the time series changes in the 1–5 Hz frequency band, essentially  
 96 showing where the frequency content of the seismogram changes appreciably. Though the MERMAIDs record and  
 97 store data with a 40 Hz sampling frequency, transmission is normally done at 20 Hz to save transmission time and  
 98 cost, which has proven sufficient for accurate picking.

### 99 3 Experiment 1

100 Picking for both experiments is done for clusters of closely located events, arranged in order of decreasing magni-  
 101 tude. This allows for the analysts to get used to the peculiarities of data coming from certain regions while learning  
 102 to pick data with a high signal-to-noise ratio before continuing on to lower magnitudes. Only events with at least one  
 103 MERMAID pick are included in our data set. For the duplicate picks of experiment 1 we select six clusters of events

**Table 1** Events used in Experiment 1

Date	Lat	Long	Depth	$M$	$N_{MH}$	$N_{GSN}$
Cluster A						
2018/08/24	-11.035	-70.781	618.2	7.1	25	211
2019/01/05	-8.165	-71.587	580.0	6.8	72	146
2022/06/08	-9.047	-71.178	622.7	6.5	124	222
Cluster B						
2018/10/07	-28.194	-179.196	400.0	5.6	36	203
2018/11/29	-27.361	-178.061	256.5	5.1	9	111
2021/09/22	-27.556	-178.810	352.5	5.0	70	111
Cluster C						
2023/02/09	-26.649	-178.300	263.8	4.9	73	44
2020/11/22	-28.334	-179.274	396.9	4.5	24	44
2022/09/20	-27.760	-178.995	356.9	4.5	34	48
2020/02/12	-26.754	-178.361	320.5	4.1	13	27
2020/11/15	-26.568	-178.157	233.3	4.1	6	40
2019/09/01	-27.241	-178.368	322.8	4.0	5	1
Cluster D						
2020/01/28	19.350	-78.847	10.0	7.7	88	89
2021/08/14	18.434	-73.482	10.0	7.2	47	194
2020/01/07	17.824	-66.823	13.7	6.4	18	108
Cluster E						
2022/11/22	-9.820	159.603	14.0	7.0	127	183
2022/11/22	-9.820	159.459	10.0	6.0	87	151
2021/10/15	-8.878	158.464	33.0	6.4	17	69
Cluster F						
2020/03/14	-27.695	-175.697	15.0	6.4	108	217
2021/08/14	-22.421	-174.552	10.0	5.6	44	86
2021/06/26	-28.330	-176.549	10.0	5.3	63	39
2021/04/16	-30.414	-177.766	10.0	5.0	22	39
2021/06/03	-24.984	-175.696	10.0	4.8	32	7
2021/04/17	-27.192	-175.923	10.0	4.4	28	11

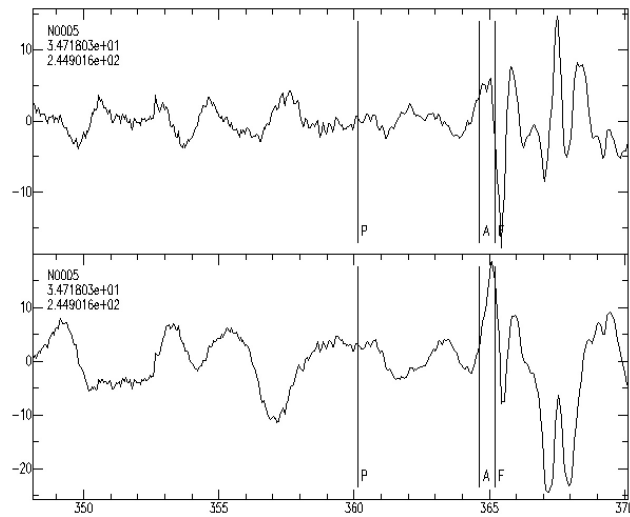
104 listed in Table 1. The last three columns in this table list the magnitude  $M$  (which is the moment magnitude  $M_w$  when  
 105 available), the number of picks from MERMAIDS ( $N_{MH}$ ), and those from surface stations ( $N_{GSN}$ ).

106 Each event is picked by up to 12 analysts. For each event, we calculate the average pick time for each station  
 107 as well as the deviation  $\Delta t$  for each pick. The distribution of these residuals  $\Delta t$  is used as a proxy for the picking  
 108 errors. For each of the six clusters we compute the RSDR or Robust Standard Deviation of the Residuals (Motulsky  
 109 and Brown, 2006), which essentially defines the 68% confidence limit. A first RSDR estimate was used to remove a  
 110 few (26) outliers beyond 3 standard deviations before computing the final RSDR again.

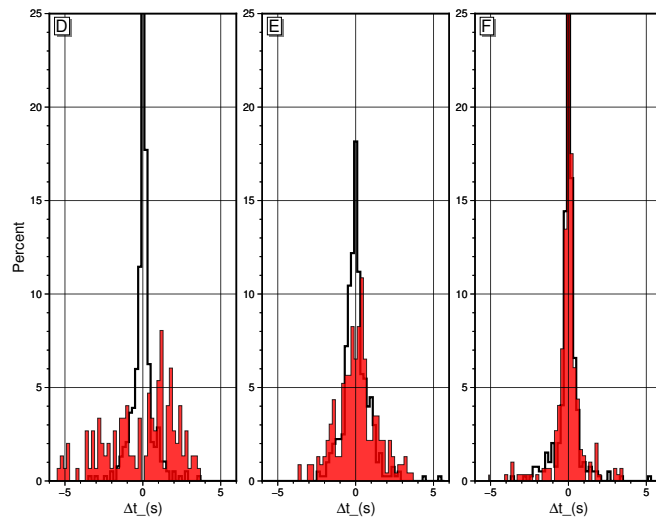
111 For the three deep clusters A,B and C, MERMAID residuals are in an acceptable range. The fact that the RSDR  
 112 for the events in the magnitude 4 range (cluster C) is smaller than that for magnitude 5 (cluster B) can probably be  
 113 explained by the fact that, even though the amplitude is smaller, the frequency of the P wave from weaker events is  
 114 higher. Also, such weak events are only observed at close or regional distances, again favouring a higher frequency,  
 115 which is easier to pick.

116 But the failure of MERMAID picks for shallow events in clusters D and E to match the precision of those from  
 117 surface stations is disappointing. Whereas the RSDR for the three deep clusters gives a distribution of  $\Delta t$  that is  
 118 comparable between MERMAIDS and surface stations, the shallow events are picked with a rather erratic distribution  
 119 of residuals, in contrast to that for the land stations (Figure 8). The exception is cluster F, which has shallow events  
 120 close to the network of MERMAIDS, resulting in easily observable high frequency onsets.





**Figure 7** SAC plot used for picking of the seismogram of 2022/06/08 recorded by MERMAID N0005. The original seismogram is at the bottom, a high-passed version (corner frequency 1 Hz) is at the top. The line indicated by P is the AK135 (Kennett et al., 1995) predicted arrival, F the AIC pick, and A the visually picked first arrival. Arrivals can be picked on either of the two plots.



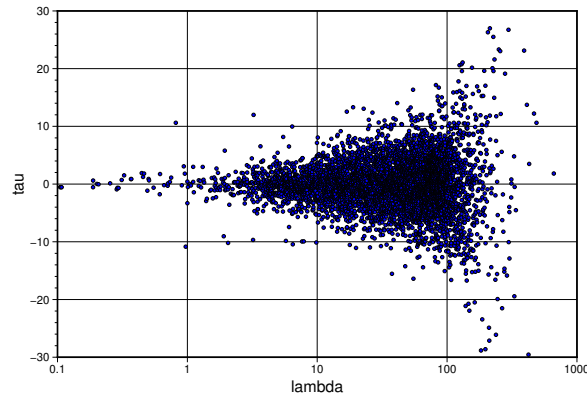
**Figure 8** Distribution of  $\Delta t$  for shallow events of clusters D, E and F observed in surface stations (thick black line) and MERMAIDs (red histogram) shows the irregular distribution of MERMAID picks for cluster D.

121 The overlap in frequency of seismic noise and that of P waves from shallow earthquakes is large, making the  
 122 identification of an onset more difficult. The failure of the events in clusters D and E to come up with a distribution  
 123 that is close to Gaussian shows that these shallow event picks are dominated by outliers. Those in the Caribbean  
 124 (cluster D) with an RSDR of 2.5 s are essentially useless for seismic delay-time tomography, where the useful signals,  
 125 i.e. traveltimes delays introduced by velocity heterogeneities, are generally smaller. Recent efforts in waveform fitting  
 126 of MERMAID seismograms by Pipatprathanporn and Simons (subm. 2024) have been successful and should signif-  
 127 icantly reduce misidentification of pP as P, which we have observed in some of our picks and suspect to be a main  
 128 cause of outliers.



**Table 2** RSDR of pick distributions (s)

Cluster	$\sigma_{MH}$	$\sigma_{GSN}$
Cluster A	0.27	0.10
Cluster B	0.48	0.24
Cluster C	0.20	0.19
Cluster D	2.50	0.40
Cluster E	1.33	0.66
Cluster F	0.39	0.29

**Figure 9** Projected data  $\tau_i$  as a function of eigenvalue  $\lambda_i$  for the most densely packed cluster of shallow earthquakes (in Tonga-Fiji). Note the paucity of small  $\lambda_i$ , indicating a high relative resolution for this subset of data.

## 4 Experiment 2

The analysis in the previous section was straightforward, since it was directly done on multiple measures of the same source-receiver path. In the second experiment we seek to confirm the findings of experiment 1 by using the interdependence of the data, as provided by the linearized tomographic equations, e.g. [Nolet \(2008\)](#):

$$Am = d, \quad (1)$$

where  $m$  is a vector of model parameters (which may include source corrections), and  $d$  are the data, scaled to unit variance. The delays  $d$  vary because the paths through the 3D Earth differ, but also because of picking errors. Whereas the delay caused by the velocity anomalies  $m$  of the Earth induces a correlation between the observed travel times because of (1), its errors are in principle uncorrelated between different source-station paths. The total picked data set available consists of 16,739 picks. Their distribution among shallow and deep events is shown in Table 3.

[Voronin et al. \(2014\)](#) project the delay time observations onto the nullspace of the matrix  $AA^T$  to annihilate the influence of the Earth's structure. If  $U$  diagonalizes  $AA^T$  then the distribution of the projected delays  $\tau = U^T d$  approaches the error distribution with variance  $\sigma_e^2$  as the eigenvalue  $\lambda_i \rightarrow 0$  since its variance satisfies:

$$\sigma_{\tau_i}^2 = \lambda_i^2 \sigma_m^2 + \sigma_e^2, \quad (2)$$

where  $\sigma_m^2$  is the variance in delays caused by heterogeneities in the Earth. If  $\lambda_i = 1$ , the signal-to-noise ratio of projected delay  $\tau_i$  is 1, but if  $\lambda_i = 0$ ,  $\tau_i$  is fully in the nullspace of  $A$  and has variance  $\sigma_e^2$ . [Nolet and van der Lee \(2022\)](#) split data into event clusters so as to reduce the size of  $A$  for each cluster while optimizing the overlap of rays (and thus the dependence of rows of  $A$ ) to obtain a large nullspace and estimate the standard errors in the ISC-EHB

**Table 3** Distribution of P wave picks among MERMAIDs and global network stations

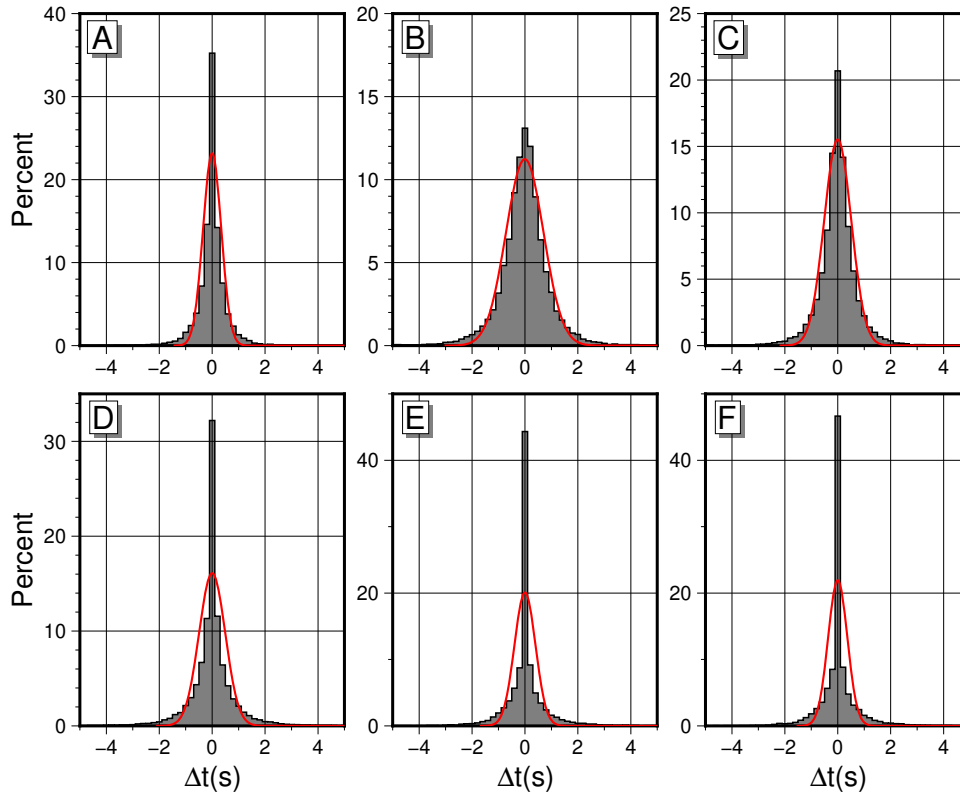
	Number of events	Total picks	MER- MAID	(is)land stations
deep (>35 km)	1147	11222	3552	7670
shallow	703	5517	1832	3685

147 catalogue of delay times.

148 We tried initially to do this also for the picks in MERMAIDs and island stations, only to find that no eigenvalue was  
 149 smaller than 0.1, even for clusters of closely spaced events, reflecting a high independence between these data caused  
 150 by the fact that the floats move around and few raypaths are therefore duplicated. Figure 9 shows the distribution of  
 151 projected data  $\tau_i$  as a function of eigenvalue  $\lambda_i$  for the most densely packed cluster of shallow earthquakes. Whereas  
 152 one clearly observes the variance decreasing with  $\lambda_i$  as predicted by (2), there is no way we can reliably estimate  $\sigma_e$   
 153 from the left part of the plot: there are few or no  $\lambda_i$  of magnitudes  $\ll 1$  for which  $\sigma_m$  can be ignored in (2) to estimate  
 154 the variance of  $\tau$  independent of model influence. While the absence of small  $\lambda_i$  is good news for any tomographic  
 155 inversion, for our experiment it means that the best we can do is to establish some lower bound for the picking errors  
 156 by investigating the a posteriori misfit to the observed traveltimes to those predicted by the tomographic model ( $Am$ ),  
 157 after imposing a reasonable regularization on  $m$  to force (1) to be overdetermined.

158 To avoid that the model parametrization introduces limitations in the resolution that contribute to the misfit we  
 159 must use a very fine grid of model voxels. We use the cubic Earth parametrization of [Charl ty et al. \(2013\)](#) which has  
 160 3,637,248 voxels to model crust and mantle. The average voxel size is 72 km at the surface, and 66 km at the bottom of  
 161 the upper mantle. Voxel thickness is adapted to fit major discontinuities but is 78 km on average. Regularization is  
 162 done by penalizing a sum of  $|m|$  and  $|\nabla^2 m|$  with  $m$  weighted by prior uncertainty – see [Nolet \(2008\)](#). For the velocity  
 163 anomaly  $\delta V_P$  we used a prior model parameter uncertainty of 1%. When additionally including source corrections  
 164 we used a prior parameter uncertainty  $\delta T_0$  of 1 s for the origin time, and 20 km for the uncertainty  $\delta h$  in depth,  
 165 longitude and latitude. [Weston et al. \(2018\)](#) give an average bias of 11 km for the ISC-EHB hypocentres, but the bias  
 166 in subduction zones – where many of the earthquakes in this study are located – is known to be much larger ([Herrin  
 167 and Taggart, 1968](#)). Regularization limits how much of the data error can ‘creep’ into the model solution to reduce  
 168 the a posteriori misfit, but cannot exclude the possibility that at least some of the data may have been erroneously  
 169 over-fitted by  $m$ . The a posteriori misfit for  $N$  univariant traveltime data defined as  $\chi^2/N = |d - Am|^2/N$  therefore  
 170 only provides a lower bound for the actual data errors. As in experiment 1, we removed outliers with a misfit beyond  
 171  $3\sigma$  after an initial, only slightly damped inversion, before calculating a final  $\chi^2/N$  estimate (our tomography code  
 172 computes the standard deviation  $\sigma$  of the misfit in the classic way, which approaches the more robust RSDR as the  
 173 distribution approaches the Gaussian).

174 Since there is ample freedom to choose the regularization, we present six tests, summarized in Table 4. The first  
 175 three (A-C) are done with (1) including corrections for the origin time and the hypocentre. These corrections are  
 176 omitted in the last three tests (D-F), as indicated by zero prior uncertainties  $\delta T_0$  and  $\delta h$  in the table. Since source  
 177 corrections require a decent azimuthal coverage of the observations, we supplement our picks with a selection of  
 178 data from the ISC catalog (until 2020) and NEIC (after 2020). We divide the source azimuth into six sectors of  $60^\circ$  and  
 179 require the combined data set to have at least four azimuth sectors with two or more data. A small number of events



**Figure 10** Histograms of the a posteriori fit of delay times for Tests A-F. The red line shows a Gaussian distribution with the RSDR as standard deviation.

**Table 4** Results of inverting (1) with different regularizations.

Test	$\delta T_0$ (s)	$\delta h$ (km)	data	$\chi^2/N$	$\delta V_p^{loc}$ (%)	outliers (%)	$\delta V_p^{max}$ (%)	depth (km)	RSDR (s)
A	1	20	picks only	1.05	-6.1	0.1	10.6	68	0.34
B	1	20	picks + cat	1.82	-5.6	0.8	9.3	34	0.71
C	1	20	picks + cat	1.01	-6.8	0.8	15.3	68	0.51
D	0	0	picks + cat	1.00	-8.5	2.7	469.8	2869	0.50
E	0	0	picks only	1.19	-10.4	3.3	101.6	11	0.38
F	0	0	picks only	1.04	-11.0	3.3	116.7	11	0.36

180 not satisfying that criterion were rejected for these tests. The added traveltime picks from the catalogs are chosen as  
 181 closely as possible to the source and such that the azimuths are as evenly distributed as possible. Results are shown  
 182 in Table 4 and Figure 10.

183 We monitor the model norm so that we can diagnose instabilities caused by data errors. However, the root mean  
 184 squared (RMS) norm of the global model  $m$  is not very useful since we are focusing on the South Pacific. Therefore,  
 185 as an indication of the model values, the table lists two proxies for the model norm:  $\delta V_p^{loc}$  is a local average  $d \ln V_p$  in  
 186 percent found between 178-180°E, 30-32°S at a depth of 68 km (the location of a large negative anomaly), and  $\delta V_p^{max}$   
 187 is the largest (absolute) anomaly in percent found throughout the whole model. The depth where this maximum is  
 188 found is listed in the next column.

189 Test A, with only our own picks, serves to check on the internal consistency of the picks. We damp to get an  
 190 overall misfit  $\chi^2/N \approx 1$ , which gives an RMS misfit of 0.63 s, close to the prior error of 0.57 s assigned to most picks,  
 191 as expected. However, the RMS estimate is heavily influenced by outliers. The RSDR, which is stable in the presence

of outliers, is 0.34 s and of the same order of magnitude as the standard deviations found in experiment 1 for deep events, but lower than for the shallow ones. Since the inversion mixes both deep and shallow events, this indicates that some of the errors are being fitted by the model, but it does not invalidate the results of experiment 1 (it would only if the RSDR exceeded those errors). The values for the model norm proxies are acceptable for a tectonically active region.

In test B we add the catalogue data to better constrain the event corrections, but leave the damping the same. As a result  $\chi^2/N$  is not close to 1 (it is 1.82), which could still be acceptable if our error estimates are in error by about 35% ( $\sqrt{1.82} = 1.35$ ). The RSDR more than doubles to 0.71 s (the RMS estimate was 1.05 s). This is still consistent with experiment 1, but it does indicate an incompatibility between catalogue data and our picks. One explanation for the increased RSDR is that catalogue data have originally been fitted with a source in the wrong location, which became incompatible after adding MERMAID data to complete the azimuth coverage. There may also be a difference in the quality for catalogue picks that were possibly obtained by an algorithm without human intervention. In the case of island station picks, there are a few duplications of catalogue data with our picks. We visually checked several of the largest discrepancies and are confident that with few exceptions our picks are accurate.

In test C we relax the damping to obtain  $\chi^2/N = 1.01$ , which lowers the RSDR to 0.51 s, but raises  $\delta V_p^{max}$  to 15.3%, a clear indication that model variations are trying to compensate for inadequate source corrections.

To further investigate the role of the hypocentre in the a posteriori misfit, we eliminate the source corrections in tests D-F. The source location and time is thus tailored to the catalogue data and ignores the new information from MERMAIDS. Inverting the combined data set (test D) while reducing the damping such that  $\chi^2/N \approx 1$  results in a severe instability, with the model parameters exceeding 100% outside the region of interest. Using the same damping as in (A) for picks only, still gives unacceptably large  $\delta V_p^{max}$ . We conclude that source corrections are absolutely necessary, since in this case the model velocity anomalies are trying to compensate for the absence of source corrections. Changing the damping to obtain a fit near 1 does not change that conclusion (test F).

## 5 Conclusions

Even though MERMAIDS operate in a noisy oceanic environment, the onset of P waves can be picked with an accuracy well below 0.5s if the earthquake is located below the crust. For crustal earthquakes the accuracy varies strongly with the frequency content of the P wave. Discrepancies show up when our arrival times are inverted together with those published in catalogue data, which points to the significance of event mislocations in oceanic areas. Such mislocations can be avoided by employing MERMAIDS in oceanic areas of interest, such that the azimuthal coverage is improved. The source corrections themselves are obviously of interest, and since the dominant drift of the floats is westwards, more data on them is steadily accumulating. We shall study them in a follow-up paper.

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## Data and code availability

The MERMAID metadata and seismograms are available at the EarthScope Consortium data center (<https://www.earthscope.org>, formerly Incorporated Research Institutions for Seismology [IRIS]) under the network code “MH” (doi: 10.7914/S-N/MH). With a few exceptions (P0006, P0007, P0008, P0010, and P0016) these seismograms are embargoed for two years after acquisition. ISC-EHB data (Weston et al., 2018; Engdahl et al., 2020; International Seismological Centre, 2020) are available from <http://www.isc.ac.uk/isc-ehb/>, and NEIC arrival times from <https://www.sciencebase.gov/catalog/item/5d110ca0e4b0941bde550412> All websites were last accessed on Nov 8, 2024.

## Competing interests

The authors have no competing interests.

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